Overview on Data Assimilation Activities in COSMO





influence of biases on COSMO-DE: bias corrections:



 'new PBL': greatly improves diurnal cycle of convective precip except for 12-UTC runs, leads to significant warm bias (of forecasts) in low troposphere at noon i.e. COSMO-DE needs excessive instability on a large scale to produce sufficient convective precip on its own





- correct RS92 humidity obs (Miloshevich et al., 2009, *J. Geophys. Res.*)
- temperature, surface pressure: test L65 with 'near-old PBL'





PP KENDA (Km-scale ENsemble-based Data Assimilation)



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- Task 1: General issues in the convective scale (e.g. non-Gaussianity)
 - obs f.g. statistics: find variable transformations to get near Gaussianity \rightarrow COSMO Met Services focus on LETKF, cease on SIR
 - − COSMO-DE EPS: fixed perturbations of model physics param. → non-Gauss. → in LETKF, need more Gaussian perturbations → stochastic physics
- Task 2: Technical implementation of an ensemble DA framework / LETKF
- Task 3: Tackling major scientific issues, tuning, comparison with nudging
- Task 4: Inclusion of additional observations in LETKF
- (Task 5: Development of a Sequential Importance Resampling (SIR) filter)





LETKF (COSMO) : method



- implementation following Hunt et al., 2007
- basic idea: do analysis in the space of the ensemble perturbations
 - computationally efficient, but also restricts corrections to subspace spanned by the ensemble
 - explicit localization (doing separate analysis at every grid point, select only obs in vicinity)
 - analysis ensemble members are locally linear combinations of first guess ensemble members





LETKF (COSMO) : technical implementation



- analysis step (LETKF) outside COSMO code
 - \rightarrow ensemble of independent COSMO runs up to next analysis time
 - \rightarrow separate analysis step code, LETKF included in 3DVAR package of DWD



- basically for verification purposes, COSMO obs operators incl. quality control will be implemented in 3DVAR / LETKF environment
 - \rightarrow future: hybrid 3DVAR-EnKF approaches in principle applicable to COSMO
- standard experimentation system not yet adapted to perform LETKF (but soon)
 - \rightarrow stand-alone scripts allow only preliminary LETKF experiments up to now





- use in-situ obs (TEMP, AIREP, SYNOP) from GME analysis (sparse density)
 → near future: use set of in-situ obs with higher density
- **3-hourly** cycles (and only up to 2 days: 7 8 Aug. 2009: quiet + convective day)
 → (near) future: 1-hourly / 30-min / 15-min cycles
- lateral (and upper) boundary conditions (BC)
 - ensemble BC from COSMO-SREPS (3 * 4 members),
 - or deterministic BC from COSMO-EU
 - \rightarrow future: ens. BC from global LETKF (GME/ICON)
- ensemble size: **32** (\rightarrow near future: \sim 40)
- initial ensemble perturbations reflect global 3DVar-B
- ensemble mean verified against nudging analysis which used much more obs

COSMO-DE: $\Delta x = 2.8$ km (deep convection explicit, shallow convection param.)

domain size : ~ 1250 x 1150 km







LETKF (COSMO), example: zonal wind at lowest model level

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4Ε.

LETKF, preliminary results: spread compared to RMSE





LETKF, preliminary results: spread compared to RMSE





Exeter, 5 October 2010



time average of RMSE against obs, and spread (7 Aug, 15 UTC - 9 Aug, 0 UTC)

(ensemble BC, with vertical localisation, adaptive **R**, $\rho = 1.15$)



 \rightarrow larger differences between analysis and first guess at observation locations

 \rightarrow underdispersive







 lack of spread: (partly ?) due to model error and limited ensemble size which is not accounted for so far

to account for it: **covariance inflation**, what is needed ?

- \rightarrow multiplicative $X_b \rightarrow \rho \cdot X_b$ (adaptive (y H(x) ~ R + H^TP_bH), Li et al., QJRMS 09)
- \rightarrow additive (stochastic physics , statistical 3DVAR-B , etc.)
- model bias
- Iocalisation (multi-scale data assimilation, 2 successive LETKF steps with different obs / localisation ?)
- update frequency
- noise control: how to deal with perturbed lateral (upper ?) BC (distort implicit error covariances in filter → limit use of obs ?)
 - hydrostatic imbalance by vertical localisation
 - digital filter initialisation ?
- non-linear aspects, **convection initiation** (latent heat nudging ?)

(\rightarrow CH project submitted: OSSE, investigate / reduce spin-up ('outer loop'))

end of 2011: ('pre-operational') LETKF suite







- stochastic physics (Palmer et al., 2009) (by Lucio Torrisi, for Jan. 2011)
- estimation and modelling of model-errors (by M. Tsyrulnikov, V. Gorin (Russia), for 2 years, start in June 2010)
 - 1. develop an objective estimation technique for model (tendency) errors (not to be confused with forecast errors), and set up stochastic model (parameterisation) for model error : $\underline{e} = u \cdot M(\mathbf{x}) + \underline{e}_{add}$ involves stochastic physics ($u \cdot M(\mathbf{x})$) and additive components e_{add}

and includes multi-variate and spatio-temporal aspects

- build statistics of forecast tendencies minus observation tendencies *d* (using pairs of lagged obs (increments) at nearly the same location)
 <u>cov(*d*) = (cov(*u*) *M*(**x**_{*j*}) *M*(**x**_{*j*}) + cov(*δM*(**x**)) + cov(*e*_{add})) * *Δt*² + 2 R
 where *δM*(**x**) = error in the model tendency due to the analysis error, cov(*δM*(**x**))_estimated by sample covariance of the ens. tendencies
 Obs: p, T, u,v, q: Synop, ship, buoy; aircraft; wind profiler (?); radar (?)
 </u>
- 2. estimate parameters for model error by fitting to statistics
- 3. develop a model-error generator and embed it in the COSMO code
- 4. validate: apply MEM to COSMO forecasts and check resulting EPS spread







Statistics: 2 months in winter and in summer. T, u, v fields, 1000-100 hPa. SYNOP & TEMP obs. on COSMO-RU domain



For most cases, both conditional mean and conditional variance are present.







• radar : radial velocity and (3-D) reflectivity

observation operators (2 PhD's, started summer 2010, DWD funding).

- implement full, sophisticated observation operators
- then develop different approximations and test their validity in order to obtain sufficiently accurate and efficient operators

Particular issues for use in LETKF:

obs error variances and correlations, superobbing, thinning, localisation

- ground-based GNSS slant path delay (after 2010, N.N.)
 - implement non-local obs operator in parallel model environment
 - test and possibly compare with using GNSS data in the form of IWV or of tomographic refractivity profiles

Particular issue: localisation for (vertically and horizontally) non-local obs







- cloud information based on satellite and conventional data
 - → DWD: Eumetsat fellowship, closing date for applications: 14 October
 - derive incomplete analysis of cloud top + cloud base, using conventional obs (synop, radiosonde, ceilometer) and NWC-SAF cloud products from SEVIRI
 - use obs increments of cloud or cloud top / base height or derived humidity
 - use SEVIRI brightness temperature directly in LETKF in cloudy (+ cloud-free) conditions, in view of improving the horizontal distribution of cloud and the height of its top
 - compare approaches

Particular issues:

non-linear observation operators, non-Gaussian distribution of observation increments







thank you for your attention



