

Status Overview on PP KENDA

Km-scale ENsemble-based Data Assimilation

Lucio Torrissi - CNMCA-Italian Met. Centre
on behalf of Christoph Schraff - DWD (P.L.)

Contributions / input by:

Hendrik Reich, Andreas Rhodin, Roland Potthast, Uli Blahak, Klaus Stephan, Africa Perianez,
Michael Bender (DWD)

Annika Schomburg (DWD / Eumetsat)

Yuefei Zeng, Dorit Epperlein (KIT Karlsruhe)

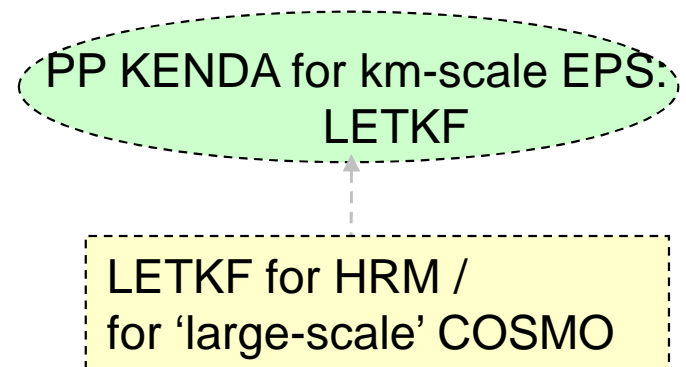
Daniel Leuenberger (MeteoSwiss)

Mikhail Tsyrlunikov, Vadim Gorin, Igor Mamay (HMC)

Lucio Torrissi (CNMCA)

Amalia Iriza (NMA)

- Task 1: General issues in the convective scale (e.g. non-Gaussianity)
- Task 2: Technical implementation of an ensemble DA framework / LETKF
- Task 3: Tackling major scientific issues, tuning, comparison with nudging
- Task 4: Inclusion of additional observations in LETKF

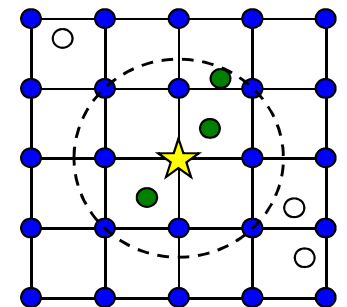


Status Overview on PP KENDA

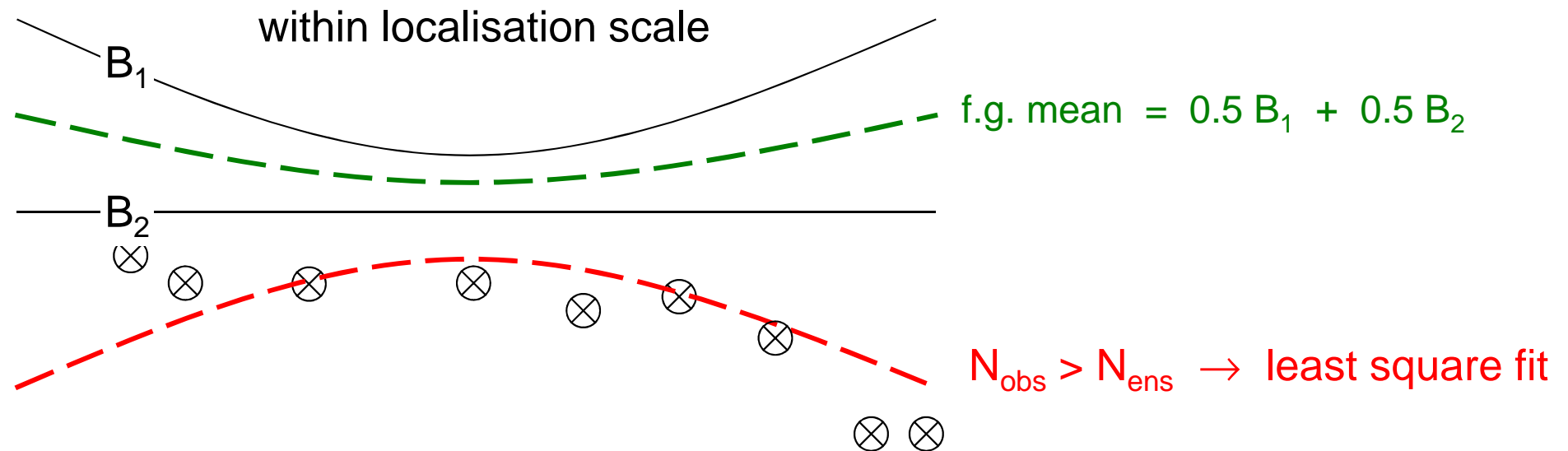
Km-scale ENsemble-based Data Assimilation

- Adaptive horizontal localization, preliminary results
- Stochastic physics implementation in COSMO model
- Use of cloud info: NWCSAF cloud products (SEVIRI/MSG)
- Work on radar observations and GNSS slant path delay

- COSMO priority project KENDA (Km-scale ENsemble-based Data Assimilation)
- implementation following Hunt et al., 2007
- basic idea: do analysis in the space of the ensemble perturbations
 - computationally efficient, but also restricts corrections to
subspace spanned by the ensemble
 - **explicit localization** (doing separate analysis at every grid point,
select only obs in vicinity)
 - analysis ensemble members are
locally **linear combinations** of first guess ensemble members



however: large N_{obs} : adaptive increase of R indicates non-optimal use of obs



→ localisation !

(or data selection / superobbing ?)

→ basic idea for adaptive localisation : keep N_{obs} constant ($> N_{\text{ens}}$, not $\gg N_{\text{ens}}$) !

LETKF, preliminary results: horizontal localisation

Caspari-Cohn function: scale $s = 100$ km

→ 0.4 at $r \cong (2)^{\frac{1}{2}} \cdot s \cong 141$ km

→ 0 at $r = 2 \cdot (10/3)^{\frac{1}{2}} \cdot s \cong 365$ km

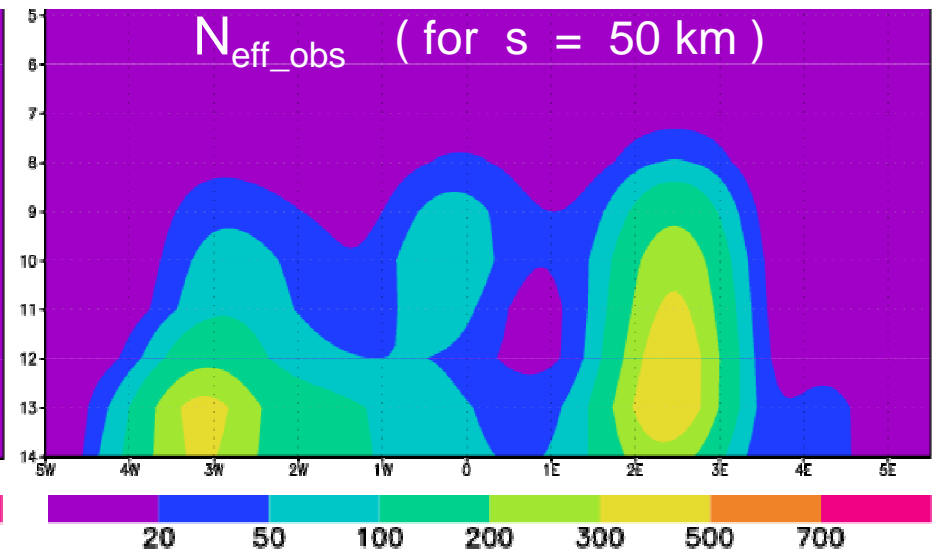
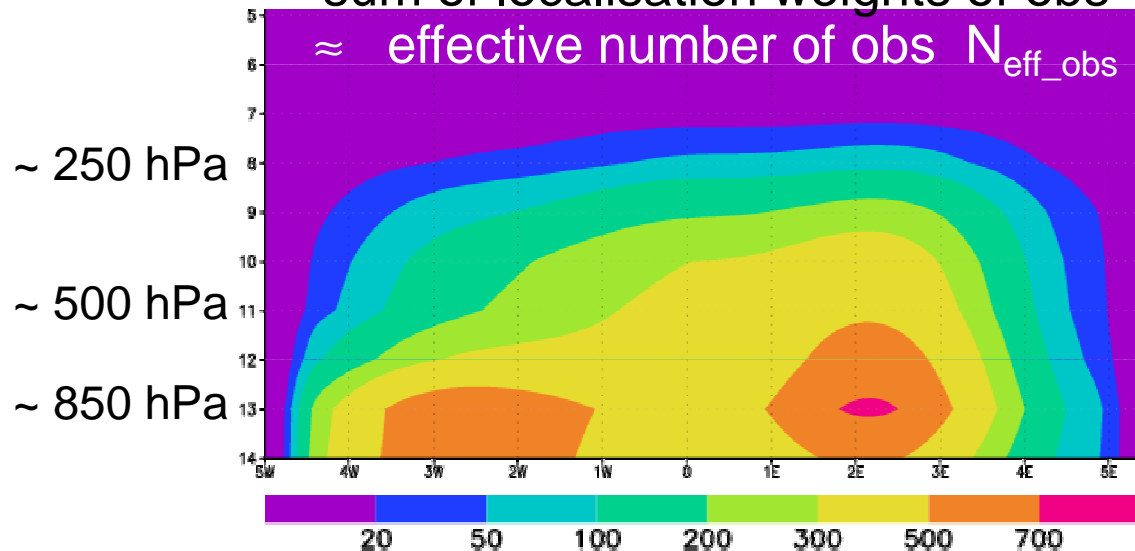
→ reduce scale : $s = 50$ km

vertical cross section

(at rot lat = 2° , 8 Aug 2009 , 12 UTC)

sum of localisation weights of obs

\approx effective number of obs $N_{\text{eff_obs}}$



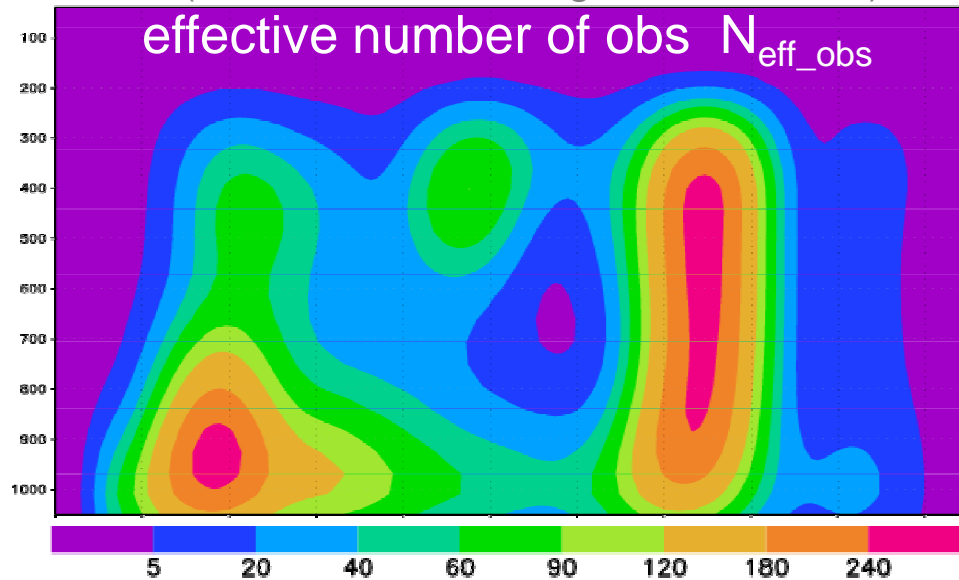
→ $N_{\text{eff_obs}} \gg N_{\text{ens}}$

→ too few degrees of freedom
in order to fit the observations

LETKF, preliminary results: adaptive horizontal localisation

Caspari-Cohn function: scale $s = 50$ km

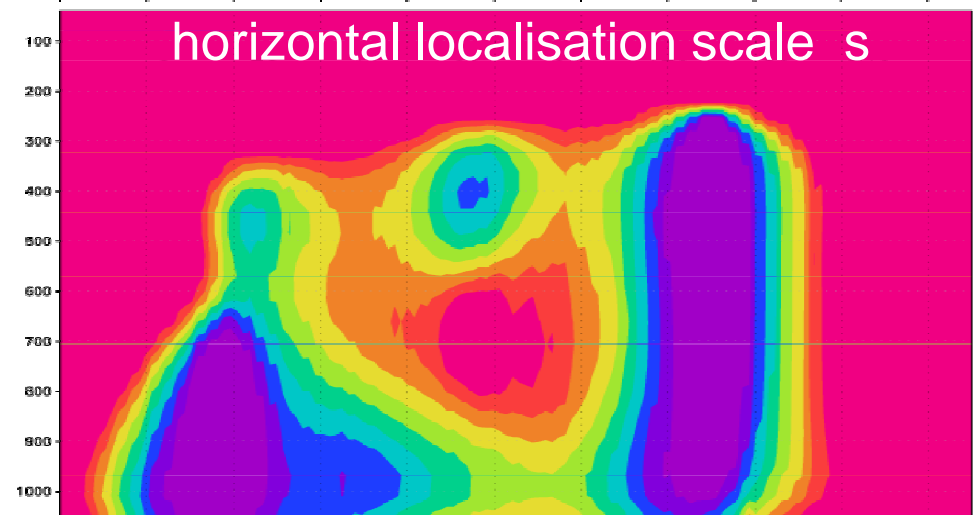
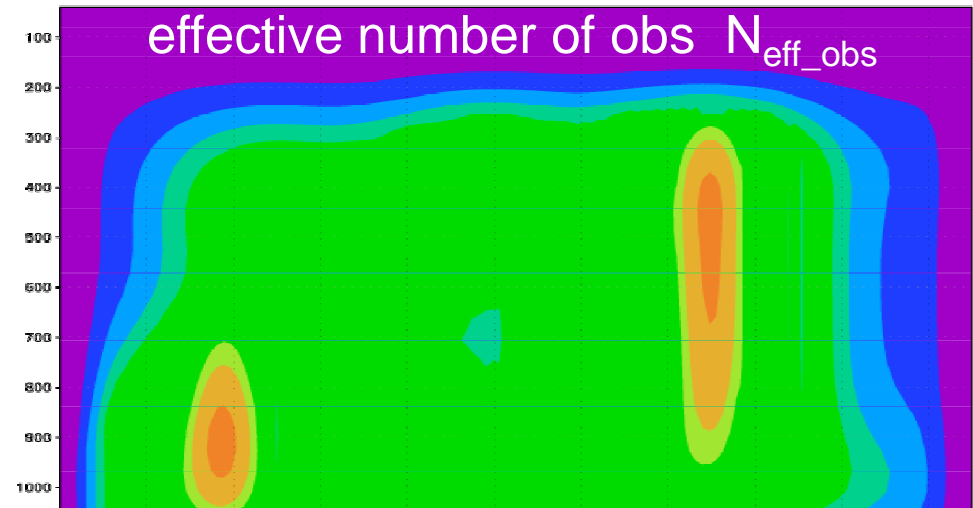
vertical cross section
(at rot lat = 2° , 8 Aug 2009, 12 UTC)



scale $s \cdot (10/3)^{1/2}$



→ adaptive scale s (32 members):
adapt s such that $N_{\text{eff_obs}} \cong 70$
and $30 \text{ km} \leq s \leq 80 \text{ km}$

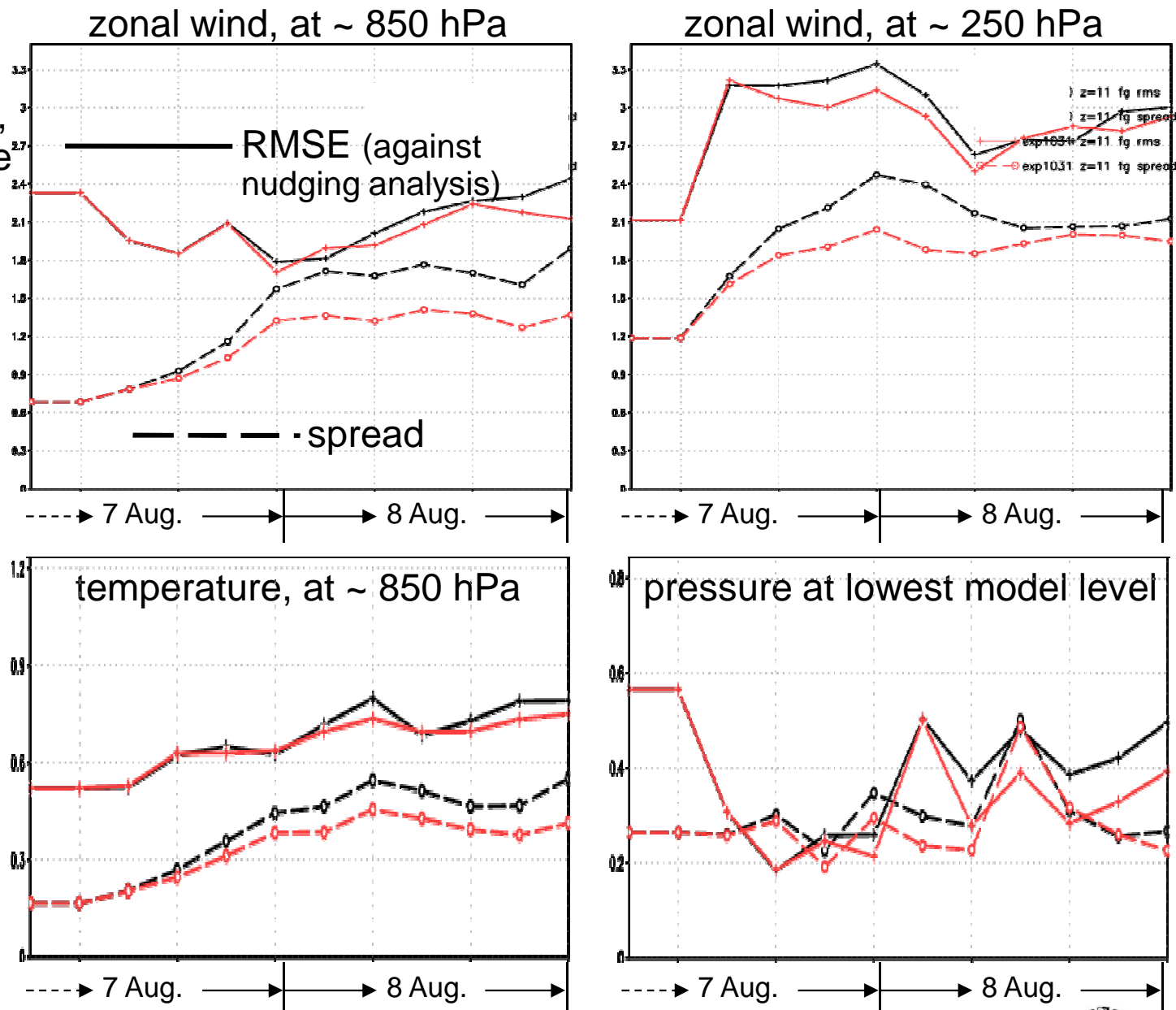


LETKF, preliminary results: adaptive horizontal localisation

first guess mean
(inner domain average)
(variable vertical localisation,
adaptive R and multiplicative
covariance inflation ρ)

$s = 50$ km
adaptive s

→ smaller spread
→ mostly smaller RMSE
(mixed results in verif
vs. upper-air obs
(T pos., wind neg.))



- new task for a pattern generator (PG)
 - purely stochastic tool to generate 4-D pseudo-random fields with selectable scales / ampl., used to generate additive perturbations / for stochastic physics (~ 0.4 FTE / y)
- stochastic physics: perturbing total physics tendency by a random factor
 - at any given grid point (Palmer et al., 2009) (Torrison)
 - modified Buizza version running
 - tuning required
 - perturb all physics tendencies in same way ?
- 2013 Ekaterina Machulskaya from SFP for (more physically based) stochastic physics ! + 1 N.N. (renewable energy project)
- additional additive inflation: - by scaled forecast differences (e.g. Bonavita et al.) ?
- 3DVAR – B ?



LETKF:

account of model error / additive inflation

Stochastic Physics

$$X_p = (1 + r \mu) X_c$$

Modified Version (in blue, differences from Buizza et al, 1999)

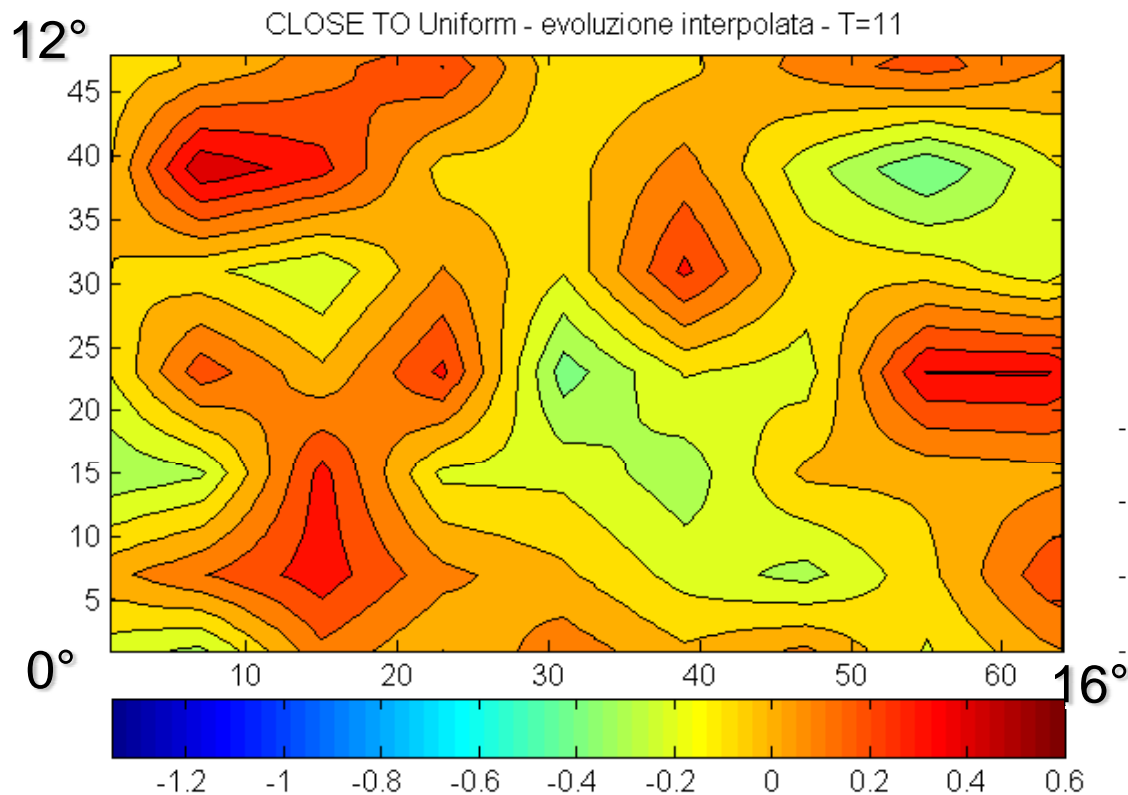
- For **all variables** (u,v,T,qv), the random numbers r are drawn from a uniform distribution in a certain range [-0.5,0.5] **or a gaussian distribution with stdv (0.1-0.5) bounded to a certain value (range= $\pm 2-3$ stdv)** on a coarse horizontal grid every n time-steps
- A tapering factor μ is used to reduce r close to the surface and in the **stratosphere** (Palmer et al, 2009)
- The perturbations of T and qv are not applied if they lead to particular humidity values (exceeding the saturation value **or negative values**)
- Spatial correlation is imposed using the same r in a whole column and drawing r for a coarse grid with spacing DL (boxes); **then they are bilinearly interpolated on the finer grid to have a smooth pattern in space**
- Temporal correlation is achieved by drawing r every n time steps (Dt); **then they are linearly interpolated for the intermediate steps to have a smooth pattern in time**
- Coarse grid SW corner is different for each member



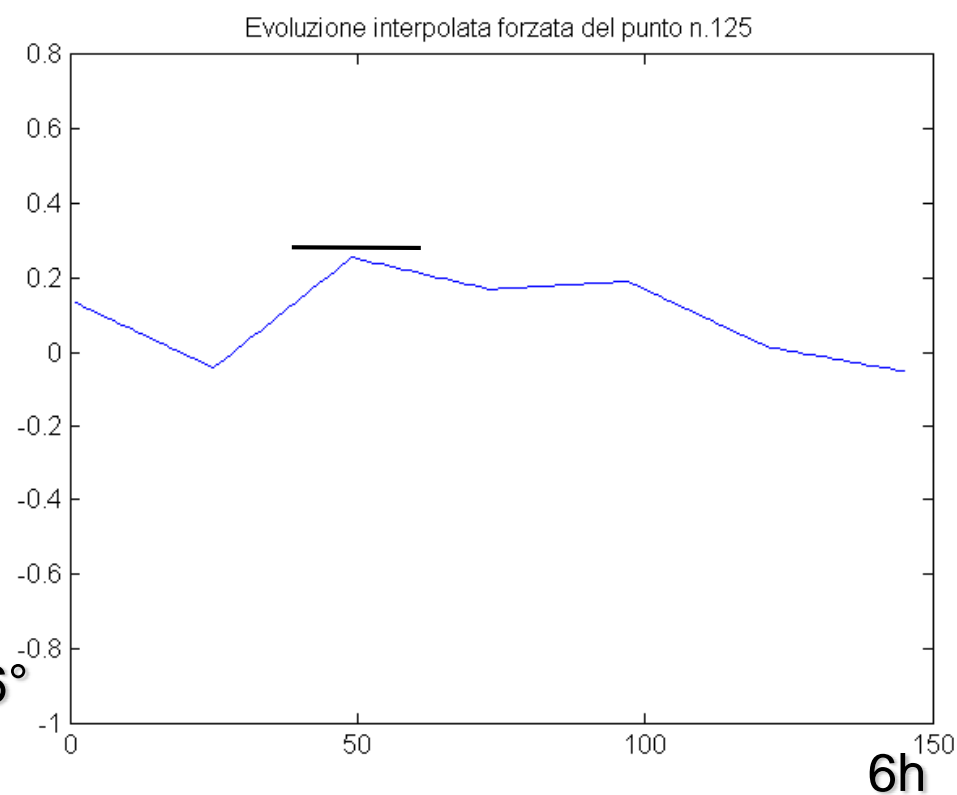
LETKF:
account of model error / additive inflation

Modified Version of Stochastic Physics

Spatial pattern



Temporal pattern



Bilin. Interpolation from coarse grid

Lin. Interpolation from coarse grid



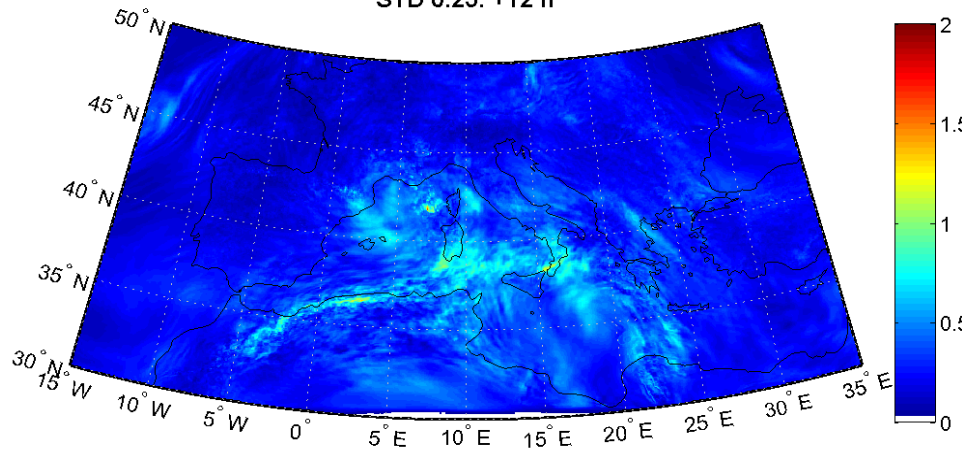
LETKF: account of model error / additive inflation

500 hPa Temperature Spread for 10 members

stdv=0.25 range=0.75

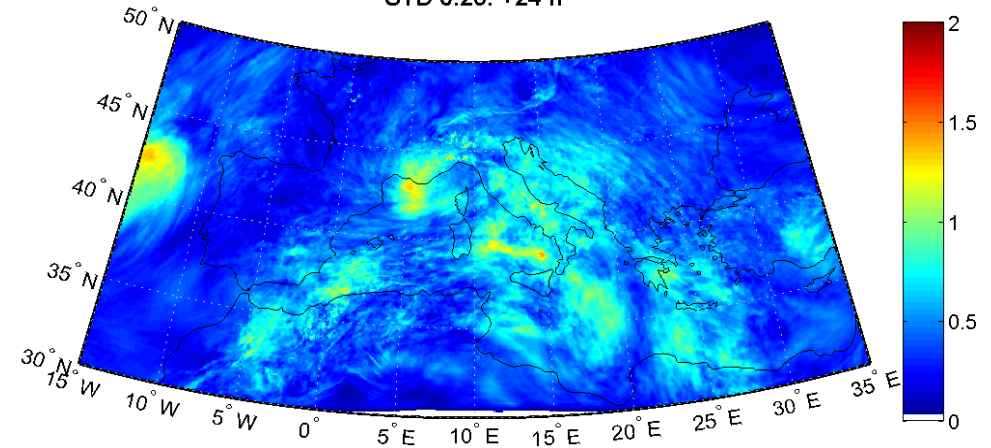
T+12h

STD 0.25: +12 h



T+24h

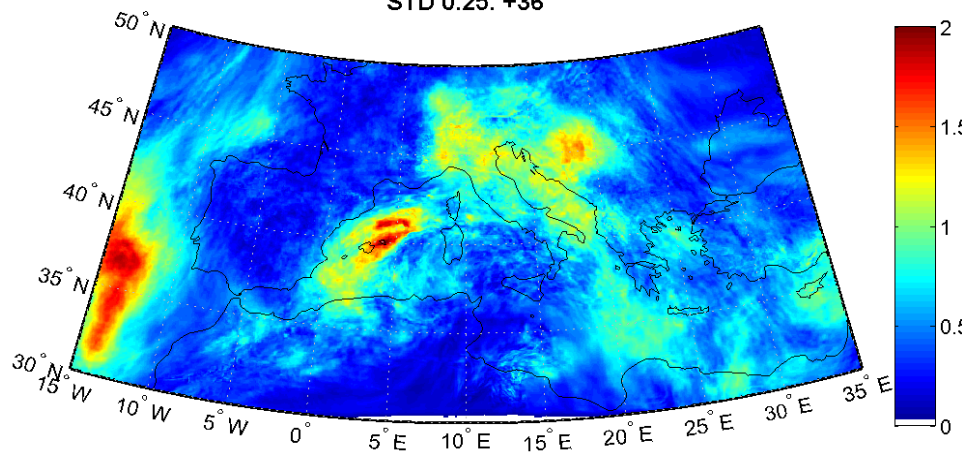
STD 0.25: +24 h



05 June 2011 00UTC

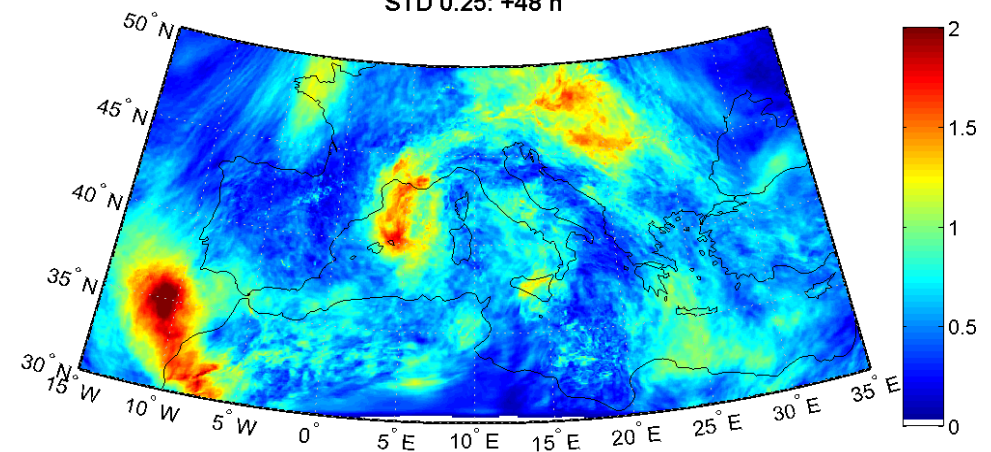
T+36h

STD 0.25: +36



T+48h

STD 0.25: +48 h

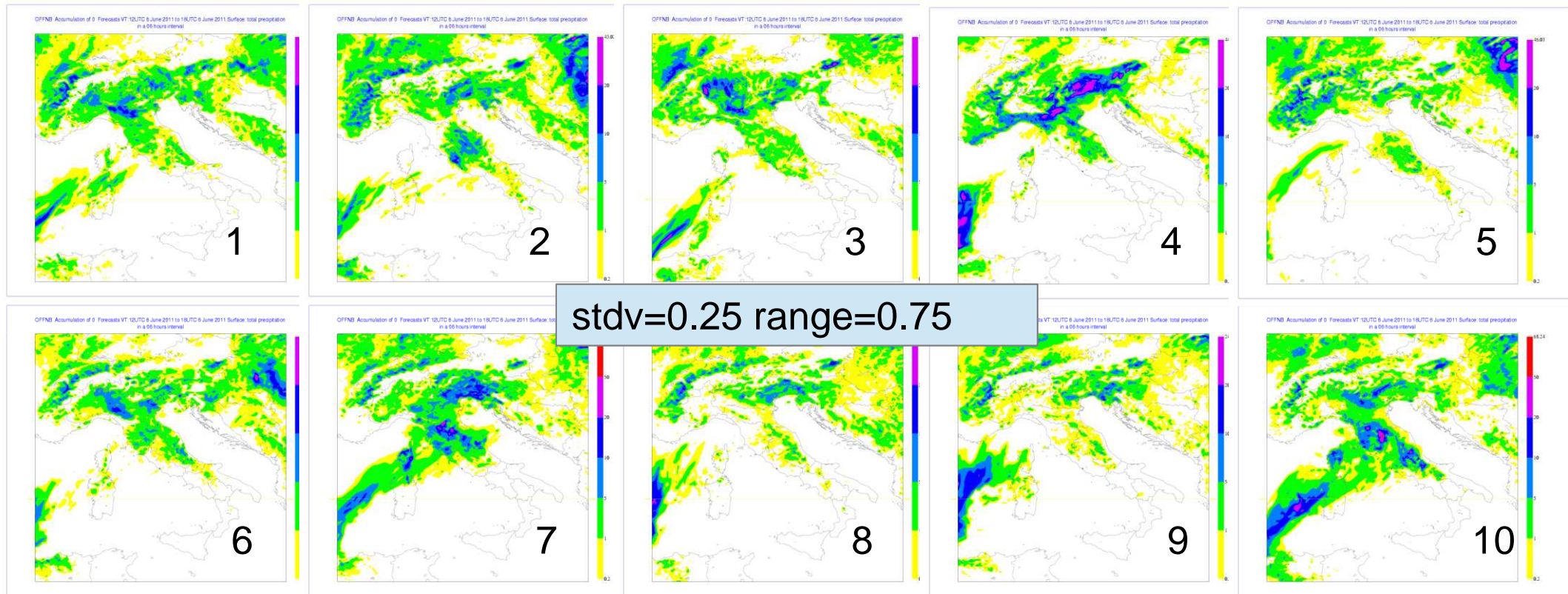




LETKF:
account of model error / additive inflation

05 June 2011 00UTC

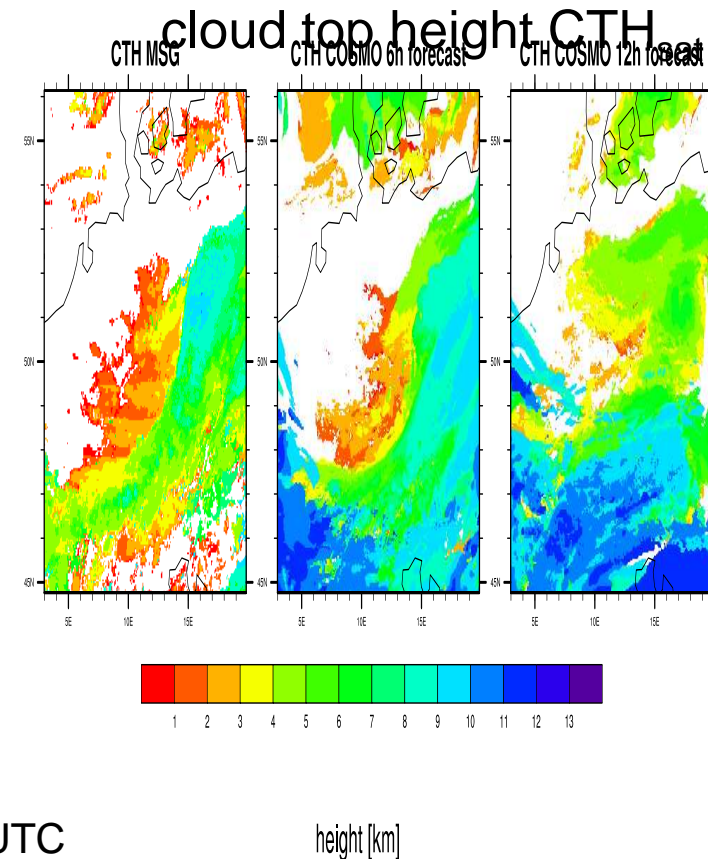
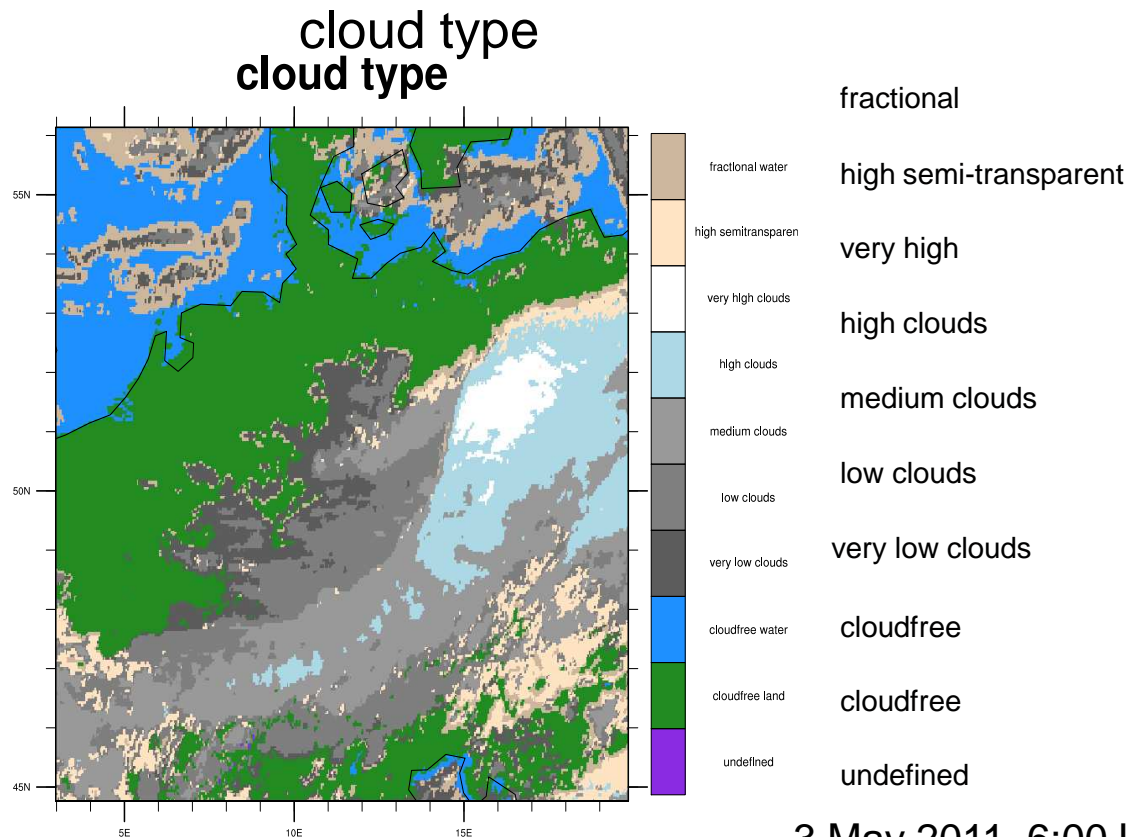
6h Accumulated Precipitation (T+36 - T+42h)



- cloud information based on satellite and conventional data
 1. derive incomplete analysis of cloud top + base height, using conventional obs (synop, radiosonde, ceilometer) and NWC-SAF cloud products from SEVIRI
use cloud top height info in LETKF
 1. (Annika Schomburg , DWD / Eumetsat)
 2. use SEVIRI brightness temperature directly in LETKF in cloudy (+ cloud-free) conditions, in view of improving the horizontal distribution of cloud and the height of its top (2013: Africa Perianez, Annika Schomburg)
- compare approaches

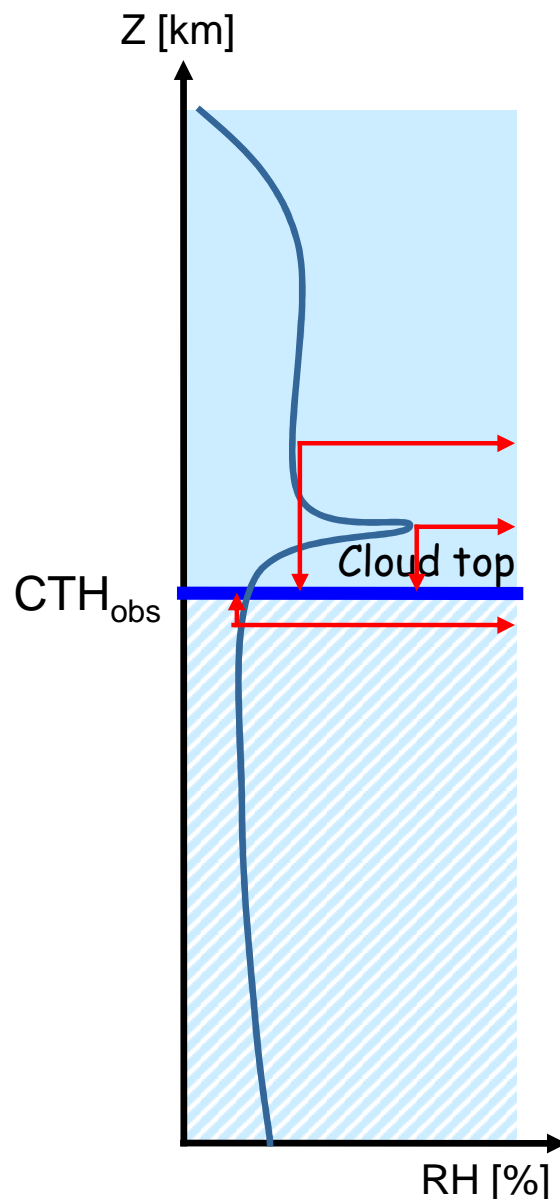
Particular issues: non-linear observation operators,
non-Gaussian distribution of observation increments

use of cloud info : NWCSAF cloud products (SEVIRI/MSG)



retrieval algorithm needs temperature and humidity profile from a **NWP model**
→ cloud top height CTH_{sat} wrong if temperature profile in NWP model wrong !
→ combine good horizontal resolution of satellite info
with good vertical resolution of radiosonde info :
use nearby radiosondes with same cloud type to correct CTH_{sat}

use of cloud info : assimilation of 'cloud analysis'



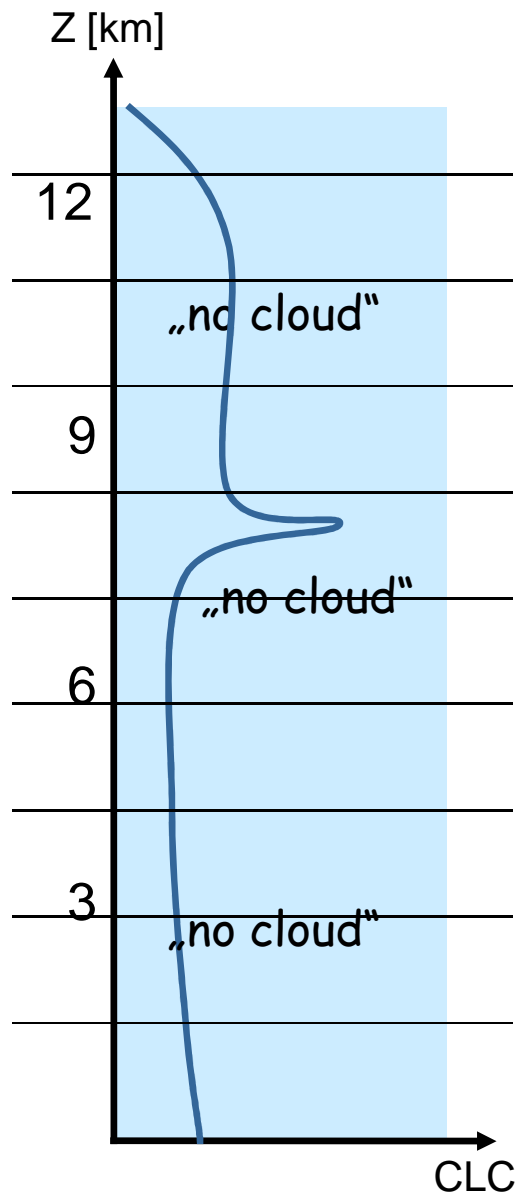
if cloud observed with cloud top height CTH_{obs} ,
what is the appropriate type of obs increment ?

- avoid too strong penalizing of members with high humidity but no cloud
 - avoid strong penalizing of members which are dry at CTH_{obs} but have a cloud or **even only high humidity** close to CTH_{obs}
- search in a vertical range Δh_{max} around CTH_{obs} for a 'best fitting' model level k , i.e. with minimum 'distance' d :

$$d = \min_k \sqrt{\underbrace{(f(RH_k) - f(RH_{obs}))^2}_{\substack{\text{function of} \\ \text{relative humidity}}} + \frac{1}{\Delta h_{max}} \underbrace{(h_k - CTH_{obs})^2}_{\substack{\text{height of} \\ \text{model level } k}}}$$

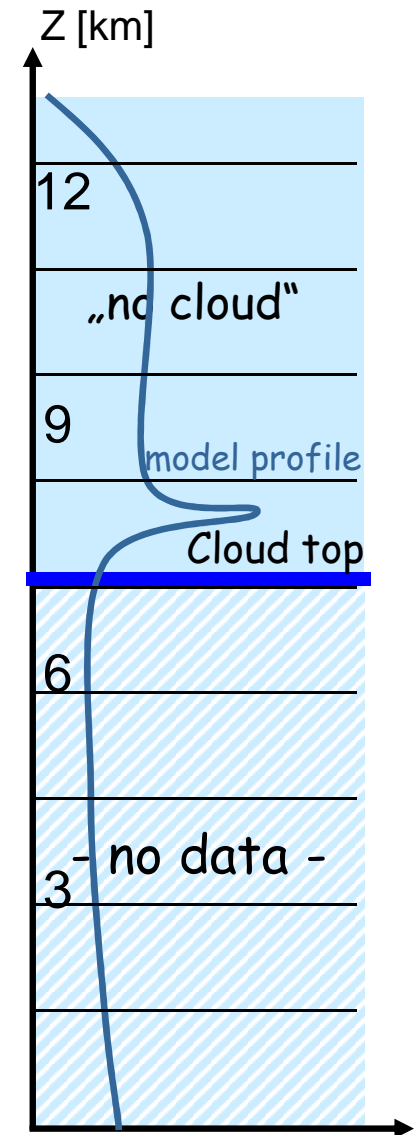
\uparrow \uparrow
 $= 1$

- use $f(RH_{obs}) - f(RH_k) = 1 - f(RH_k)$
and $CTH_{obs} - h_k$
as 2 separate obs increments in LETKF



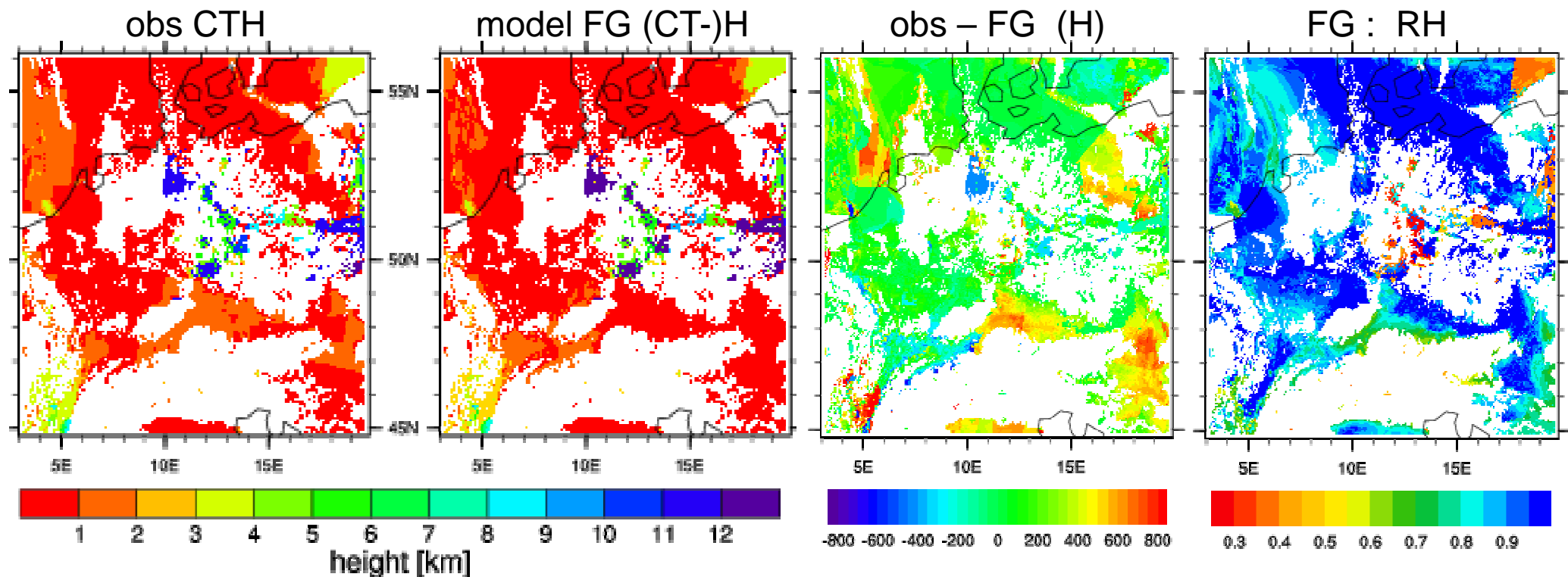
type of obs increment ,
if no cloud observed ?

- assimilate $CLC = 0$ separately for high, medium, low clouds
- model equivalent:
maximum CLC within vertical range



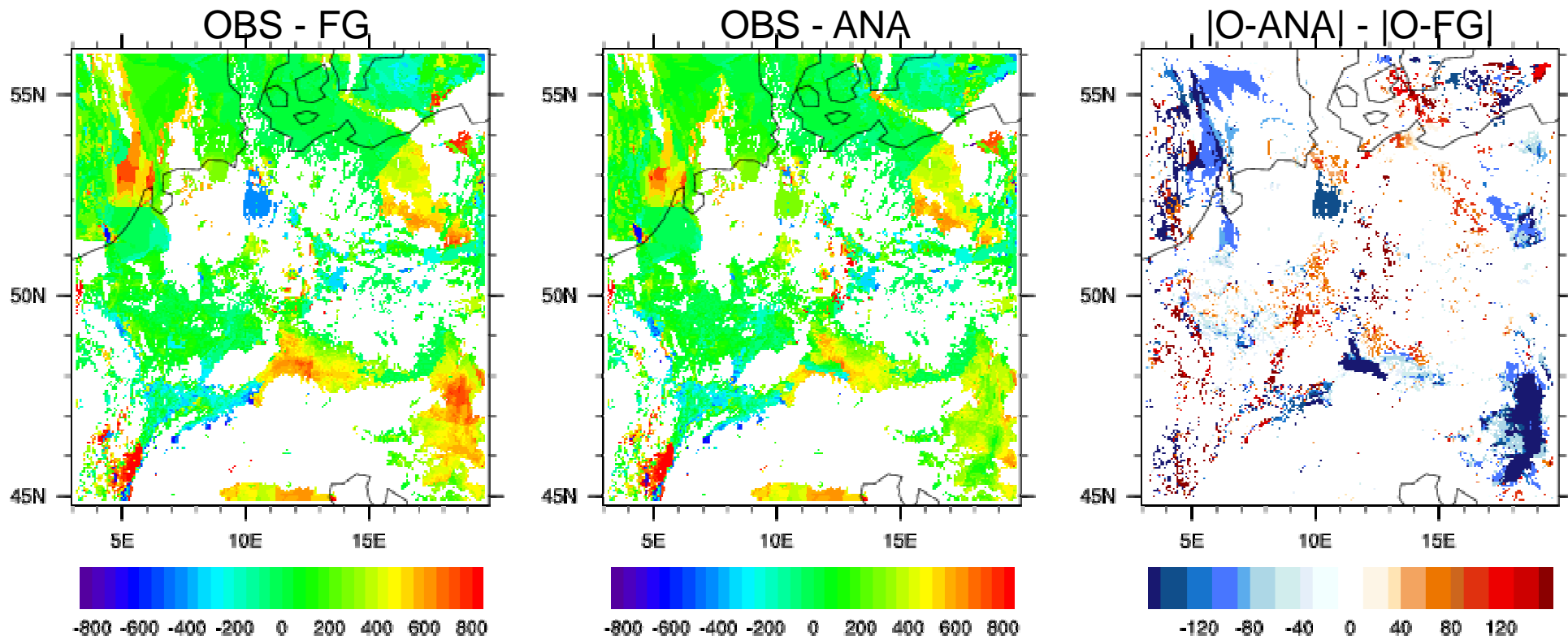
use of cloud info :
assimilation of 'cloud analysis' : example

17 Nov 2011, 6 UTC (low stratus case)
pixels where observation has clouds
(output from feedback files)



use of cloud info :
first assimilation experiment

'cloud' top height

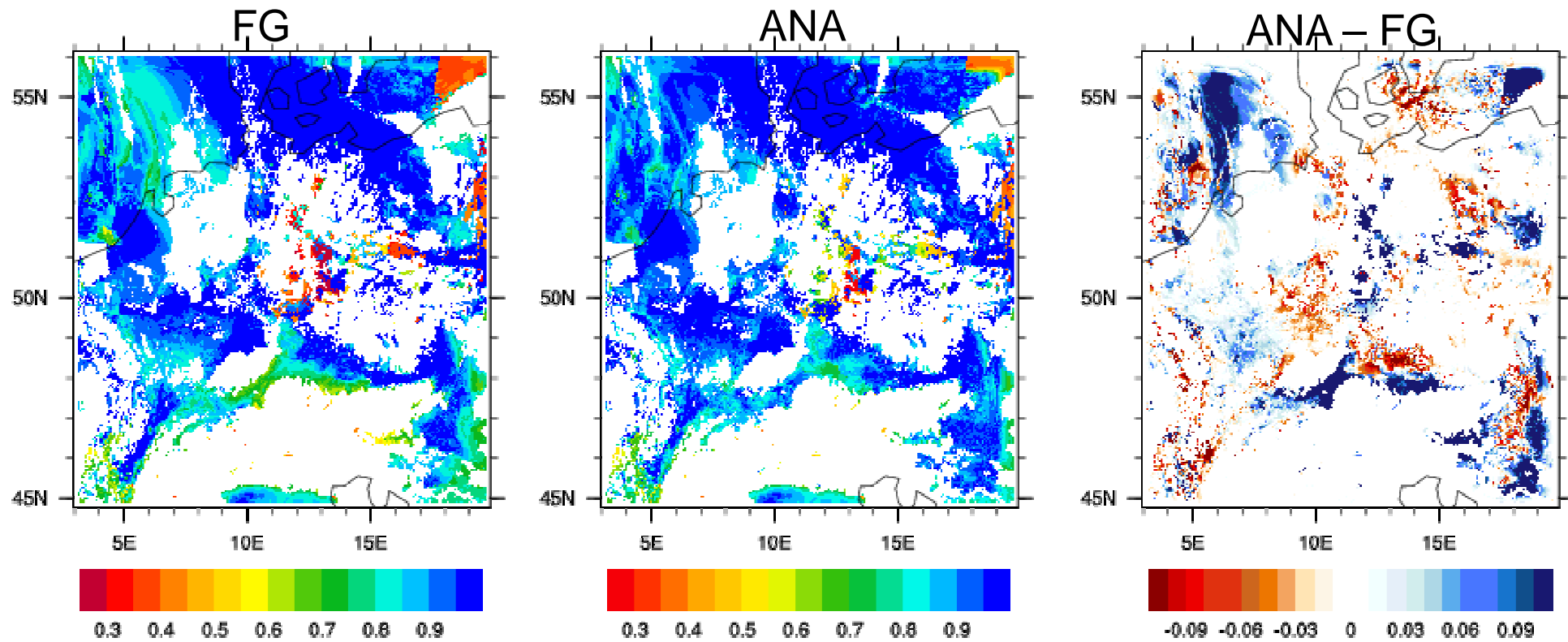


blue means ANA is better

Here: results of deterministic run in LETKF framework

(Kalman gain matrix applied to standard (unperturbed) model integration)

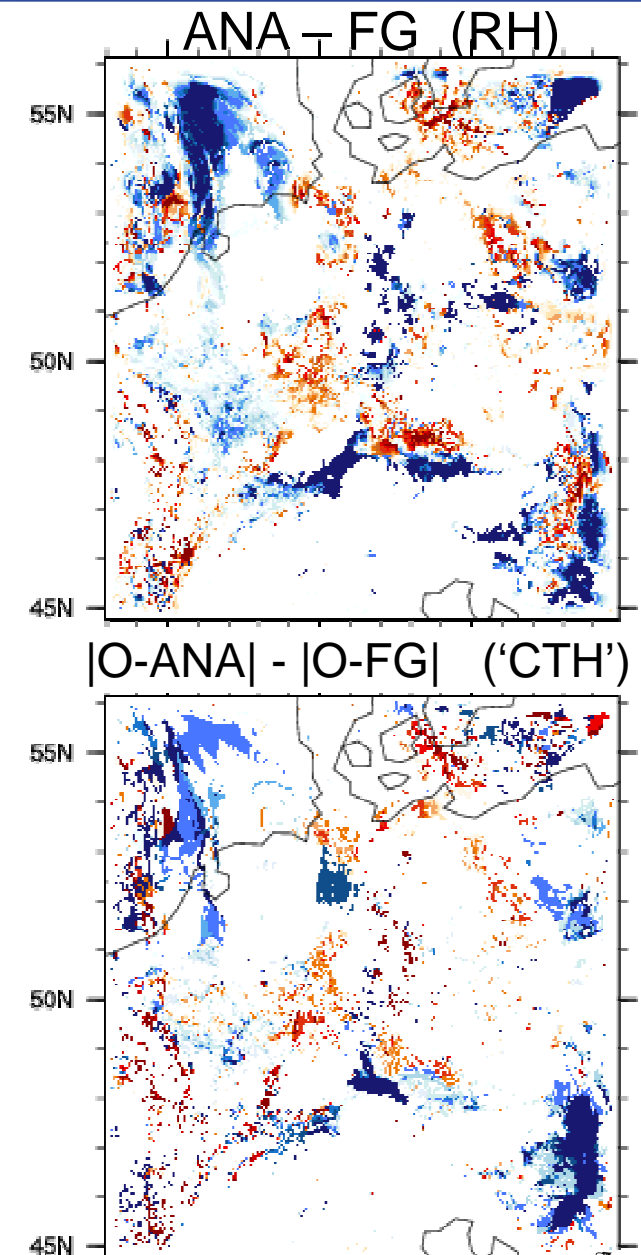
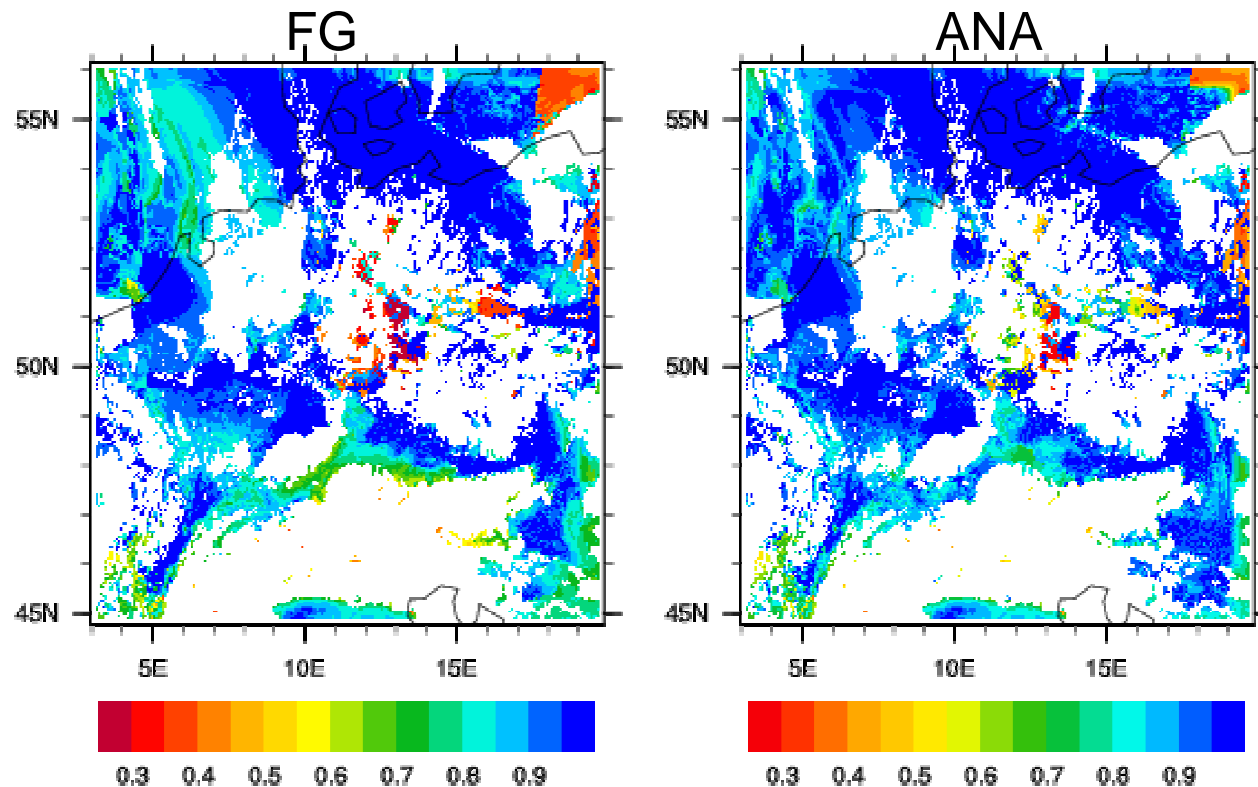
Relative humidity at 'cloud' level



Here: results of deterministic run in LETKF framework
(Kalman gain matrix applied to standard (unperturbed) model integration)

use of cloud info :
first assimilation experiment

Relative humidity at cloud level



→ LETKF draws model cloud tops closer to obs

- next :
- detailed evaluation (cross section, profiles...)
 - single observation experiments
 - tuning of obs. error, thinning, localization

KENDA: use of additional observations use of radar obs

- radar : assimilate 3-D radial velocity and 3-D reflectivity directly
 1. observation operators implemented
(Uli Blahak (DWD), Yuefei Zeng, Dorit Epperlein (PhD, KIT))
 - full, sophisticated
 - efficient (e.g. lookup tables for Mie scattering)
 - tested for sufficiently accurate and efficient approximations for DA
 2. assimilation experiments
 - technical work finished this week
 - 1 - 2 assimilation case studies (Zeng)
 - 2013: Klaus Stephan : test periods, tuning ...

Task 4.3: use of GNSS slant path delay

- ground-based GNSS slant path delay SPD
 - produce & use **tomographic** refractivity profiles (Erdem Altunac, PhD)
 - implement non-local **SPD** obs operator & use SPD (Michael Bender)
 - first implement SPD obs operator in 3DVAR package (environment for work on tomography)
 - implement simple operator (refractivity along straight line)
 - implement complex obs operator with ray tracer
 - monitoring, test e.g. impact of straight line approximation
 - then implement obs operator in COSMO (in 2013)

The 5th EnKF Workshop, Rensselaerville (New York), May 21-24, 2012

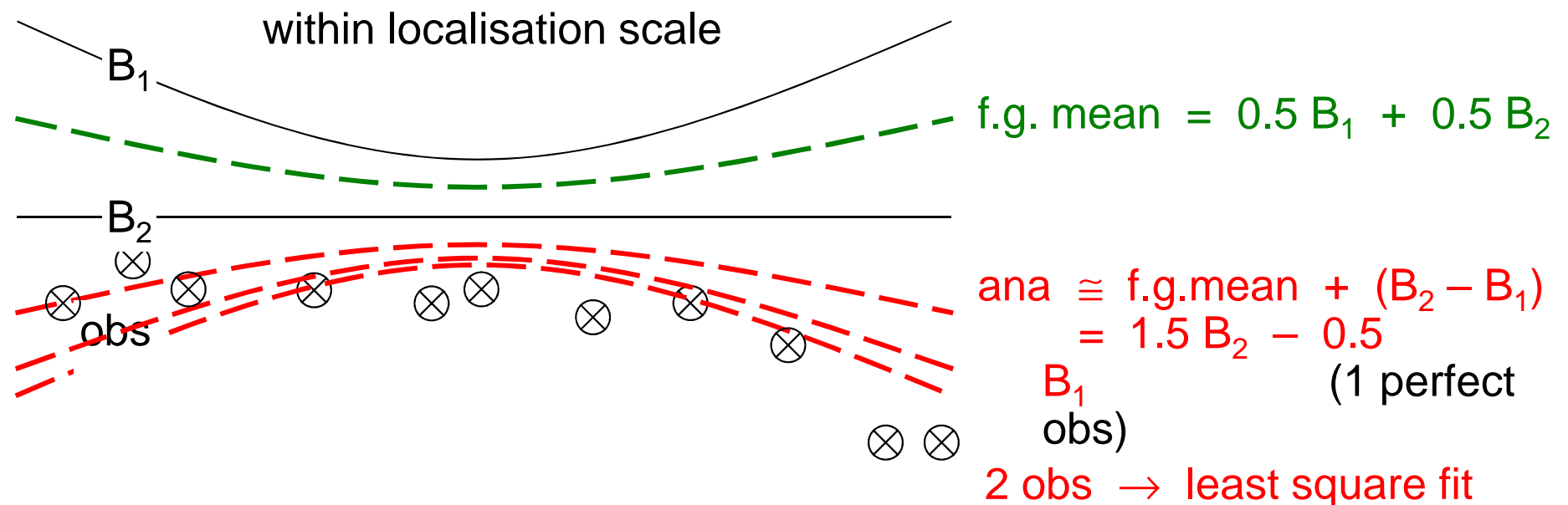
Session 5: Operational Implementations

Session chair: Lucio Torrisi (torrisi@meteoam.it)

Center	Scheme	Use	Operations	Presentation
CMC (Canada)	EnKF (Stochastic)	- EPS initialization	January 2005	Oral
NCEP (USA)	Hybrid (3DVAR/EnSRF)	- Deterministic forecast	May 2012	Oral (sess.2)
UKMO	Hybrid (4DVAR/ Local ETKF)	- Deterministic forecast - EPS Perturbations	July 2011 June 2006	Oral
MF (France)	EDA (Ensemble of 4DVAR)	- Initial B variances in 4DVAR - EPS initial perturbations together with SV	July 2008 December 2009	-
ECMWF	EDA (Ensemble of 4DVAR)	- Initial B variances in 4DVAR - EPS initial perturbations together with SV	May 2011 June 2010	Oral
CNMCA (Italy)	Regional EnKF (LETKF)	- Deterministic forecast	June 2011	Poster

thank you for your attention

- adaptive estimation of obs error covariance R
(Li, Kalnay, Miyoshi, QJRM 2009) , but our implementation: in ensemble space !



add obs, if already $N_{\text{obs}} > N_{\text{ens}}$:

- cannot be fitted well, improve analysis only slightly
- decrease analysis error !

$$\mathbf{P}_w^a = \left[(k-1) \mathbf{I} + (\mathbf{Y}^b)^T \mathbf{R}^{-1} \mathbf{Y}^b \right]^{-1}$$

\rightarrow adaptive R takes that into account and increases R

- parameterisation of model error using statistics (Tsyrlunikov, Gorin) :
 - parameterisation: $e = \mu * F_{\text{phys}}(x) + e_{\text{add}}$
 - estimate parameters by fitting to statistics from forecast tendency and observation tendency data (using a maximum likelihood based method)

failed in OSSE setup with simulated ME for finite-time 1 – 6 hr tendencies !!!

main methodological cause of failure : instantaneous ME is contaminated
in finite-time tendencies
by other tendency errors :

- trajectory drift as a result of ME themselves
 - initial errors (plus the trajectory drift due to initial errors)
- conclusion: observation accuracy and spatio-temporal coverage
far from being sufficient to reliably estimate ME !

→ task is stopped !

LETKF : implementation of verification

- tool for production of 'full' NetCDF feedback files
 - make clean interfaces to observation operators / QC in COSMO : done
 - ... integrate them into 3DVAR package : in progress
 - and extend flow control (read correct (hourly) Grib files etc.) : to be doneshould be ready by end of 2012 (for VERSUS)
- ensemble-related diagnostic + verification tool, using feedback files : (Iriza, NMA)
 - computes statistical scores for different runs ('experiments'),
 - **focus: use exactly the same observation set in each experiment !**
 - select obs according to namelist values (area, quality + status of obs, ...)
 - implementing ensemble scores (reliability, ROC, Brier Skill Score, (continuous) Ranked Probability Score)
 - main part of documentation written

- modifications in COSMO in official code (V4_24)
(e.g. in order to have a sub-hourly update frequency)
- **COSMO-DE LETKF implemented in NUMEX** and tested
(e.g. stand-alone 2-day experiment reproduced)
- GME-LETKF & ensemble INT2LM for DA cycle implemented in NUMEX,
being tested, should be available end of Sept.
→ in Oct., start first KENDA experiments in NUMEX over several days/weeks
but: - direct interpolation from 60 km to 2.8 km !
- deterministic analysis not yet implemented in NUMEX
- ensemble LBC 2013 – 2014 :
ensemble perturbations of interpolated ensemble GME fields,
added to deterministic COSMO-DE LBC

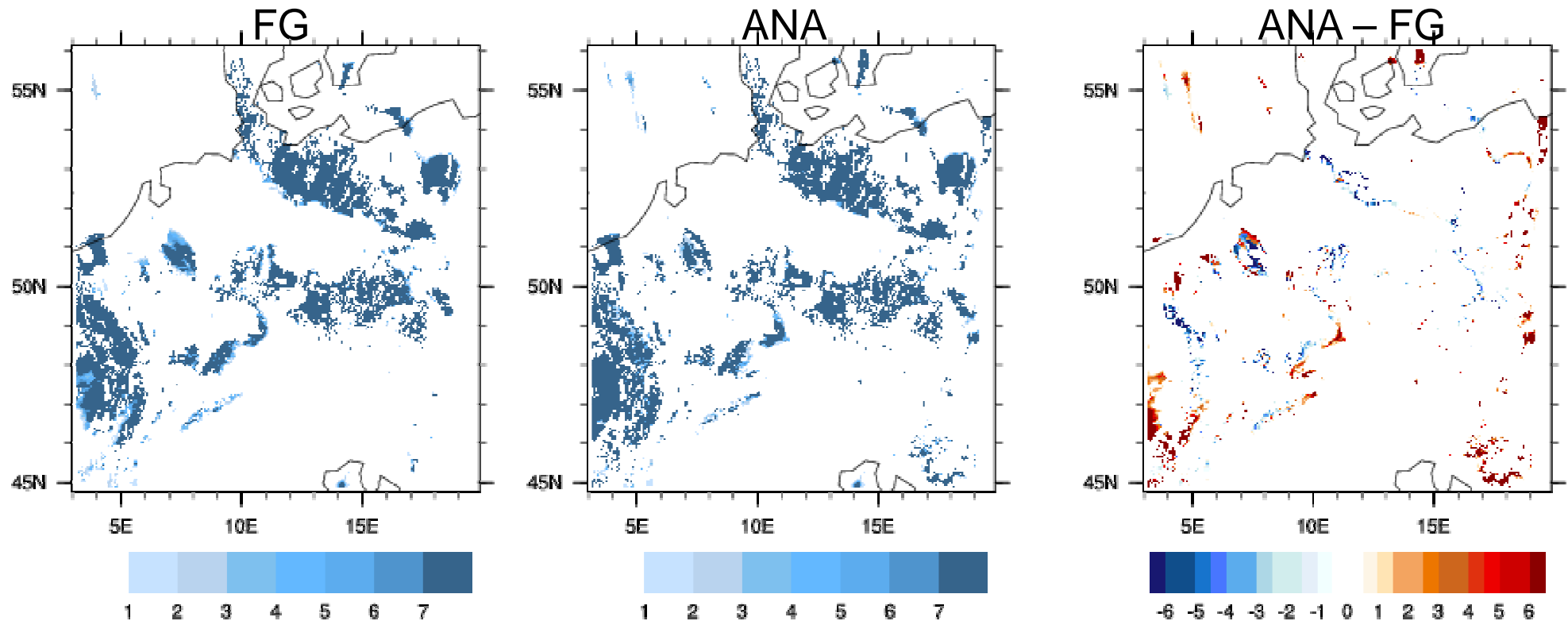
LETKF : implementation / activities

- in past year, still only preliminary LETKF experiments possible, using Hendrik's scripts:
 - up to 2 days (7 – 8 Aug. 2009: quiet + convective day)
→ 3-hourly (15-min) cycles
 - 32 ensemble members
 - perturbed LBC: COSMO-SREPS, 3 * 4 members
- therefore
 - theoretical studies, toy model experiments related to adaptive localisation
→ talk by Hendrik Reich
 - benchmark , winter school on DA , support for HErZ centre , testing (e.g. NUMEX) ...
 - only few COSMO-DE experiments related to adaptive localisation

use of cloud info : first assimilation experiment

pixels where model has clouds and observation not

Low clouds



→ from “cloudfree” obs no positive impact so far

- next steps :
- detailed evaluation (cross section, profiles...)
 - single observation experiments
 - tuning of observation error, thinning, localization

Task 3: scientific issues & refinement of LETKF

- lack of spread: (partly ?) due to model error and limited ensemble size which is not accounted for so far

to account for it: covariance inflation, what is needed ?

→ additive (see later)

→ multiplicative $X_b \rightarrow \rho \cdot X_b$

(by tuning, or) adaptive $\langle (y - H(x_b))(y - H(x_b))^T \rangle = \mathbf{R} + \rho \mathbf{H} \mathbf{P}_b \mathbf{H}^T$

→ pre-specified \mathbf{R} is used for adaptive ρ :

→ need for careful specification / tuning of obs errors

- observation error covariance \mathbf{R} :

also estimate adaptively (Li, Kalnay, Miyoshi, QJRMS 2009)

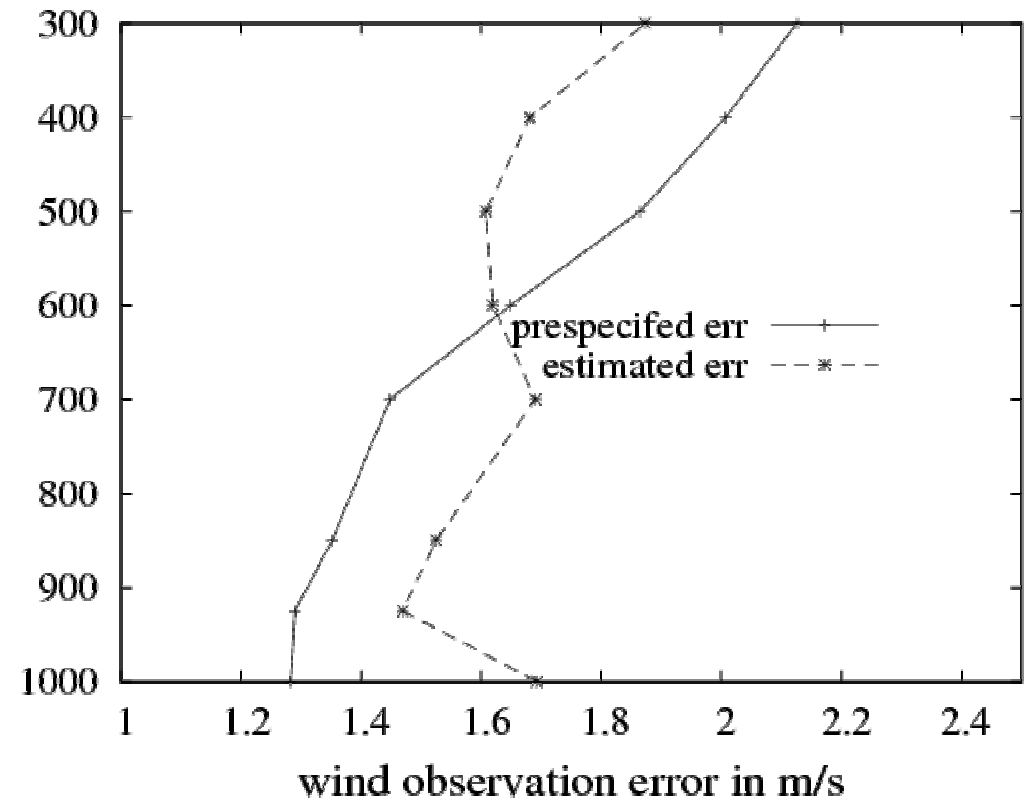
$$\langle (y - H(x_a))(y - H(x_b))^T \rangle = \mathbf{R}$$

Task 3: scientific issues & refinement of LETKF

- adaptive observation errors

$$\left\langle (y - H(x_a))(y - H(x_b))^T \right\rangle = \mathbf{R}$$

(in observation space)



- adaptive R in ensemble space:
adjusts total weight, not relative weight of obs

- localisation

