

ENSEMBLE DEVELOPMENTS IN THE COSMO CONSORTIUM

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OUTLINE

- **Operational ensembles:**
 - **COSMO-LEPS**
 - **COSMO-DE-EPS**
 - **COSMO-E**
 - **TLE-MVE (Poland)**
- **Developments**
 - **Improving the spread/skill relation**
 - **Physics perturbation**
 - **Soil/surface perturbation**
 - **Calibration and products**
 - **ICs for the ensembles**



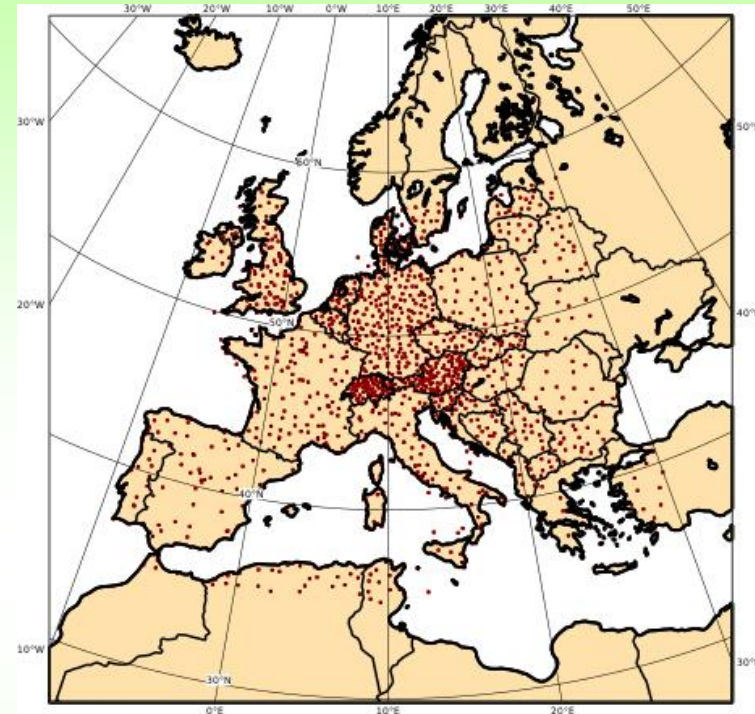
ENSEMBLES:

COSMO-LEPS



Upgrade to 20 members

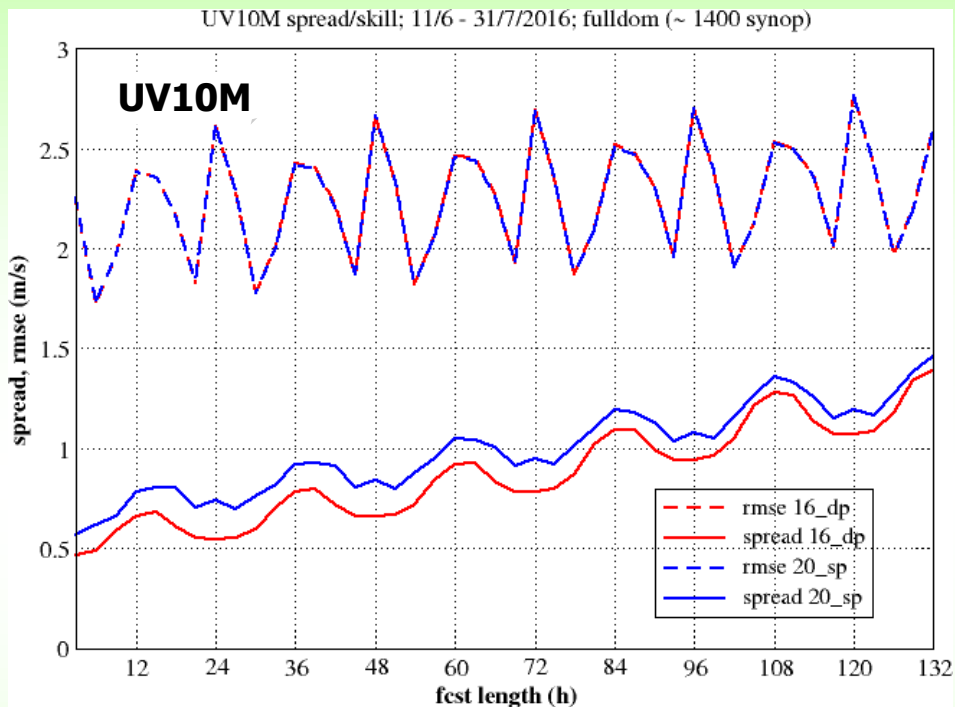
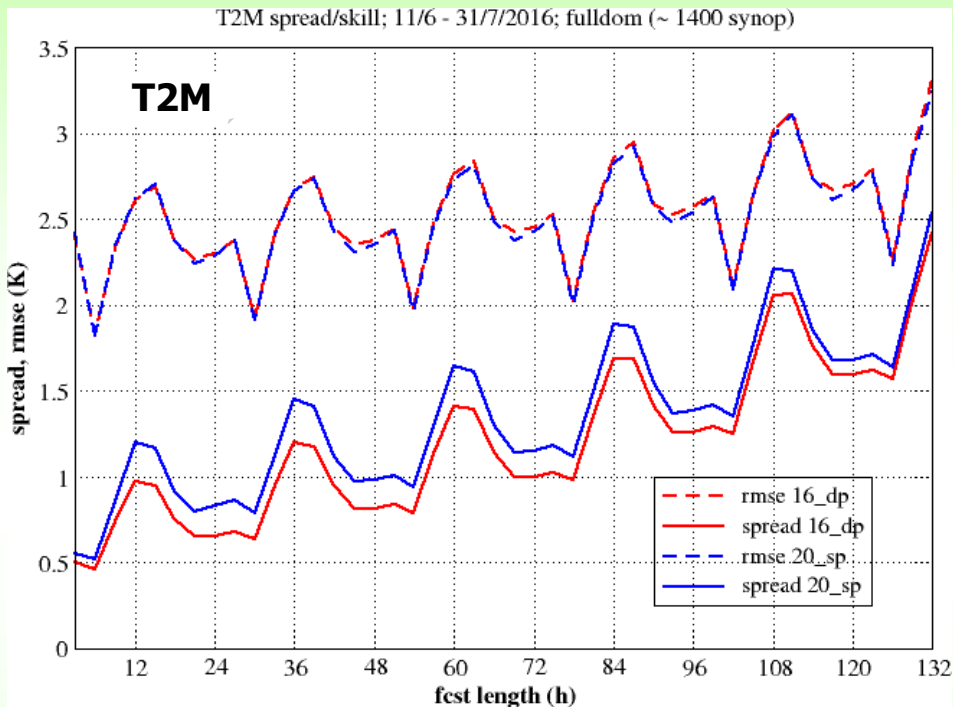
- COSMO-LEPS **current** configuration:
 - 16 members
 - COSMO with 7 km hor. res.
 - IC and BC from 16 members of ECMWF ENS (clustering)
 - parameter perturbation in physics schemes
- COSMO-LEPS experimental configuration:
 - **20 members**
 - **COSMO in single precision** (same hor. res.)
 - IC and BC from **20** members of ECMWF ENS (clustering)
 - parameter perturbation in physics schemes **+ SPPT**



Verification area (~ 1400 synop)

51 days (11/6 - 31/7/2016)

Spread/skill for T2M and UV10M

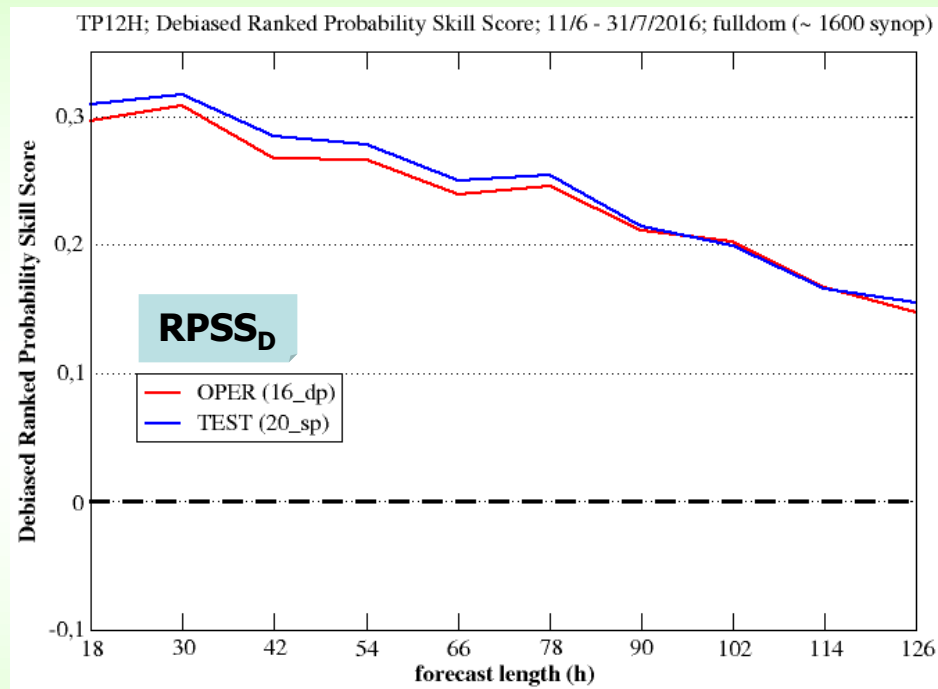
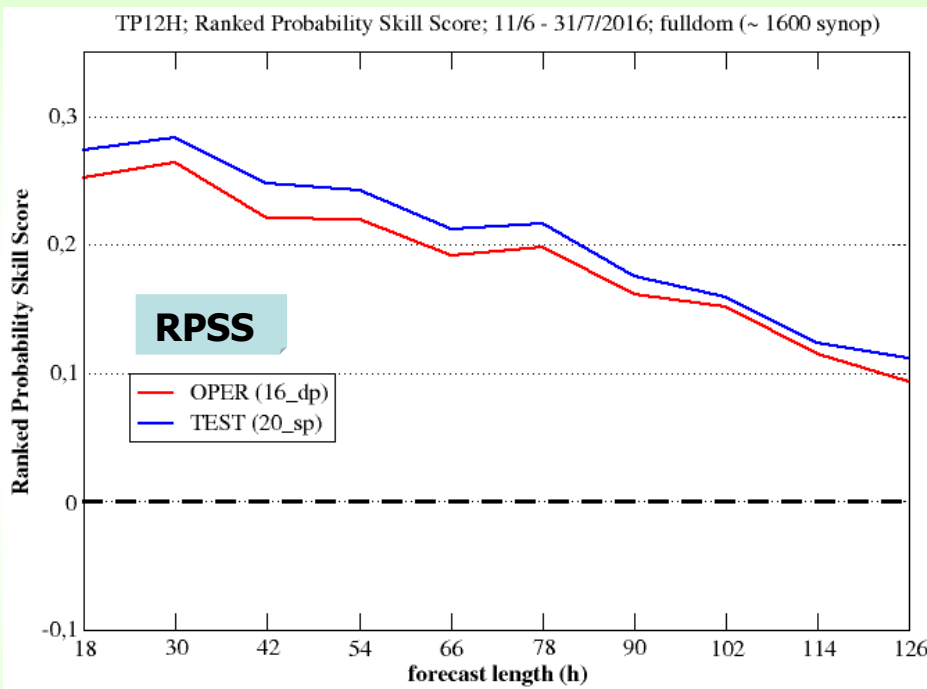


- **Larger spread for 20_sp for both variables;** in either cases, lack of spread in the short range.
- T2M: the daily cycle of the spread follows to a certain extent the cycle of the error.
- **Limited impact (if any) on the forecast skill** of the ensemble mean.

Probabilistic prediction of tp: RPSS

Consider debiased RPSS:

$$RPSS_D = 1 - (RPS / (RPS_{ref} + RPS_{ref} / N))$$



- In either cases (RPSS or RPSS_D), better performance of **20_sp** COSMO-LEPS, more evident for short ranges.

ENSEMBLES:

COSMO-DE-EPS

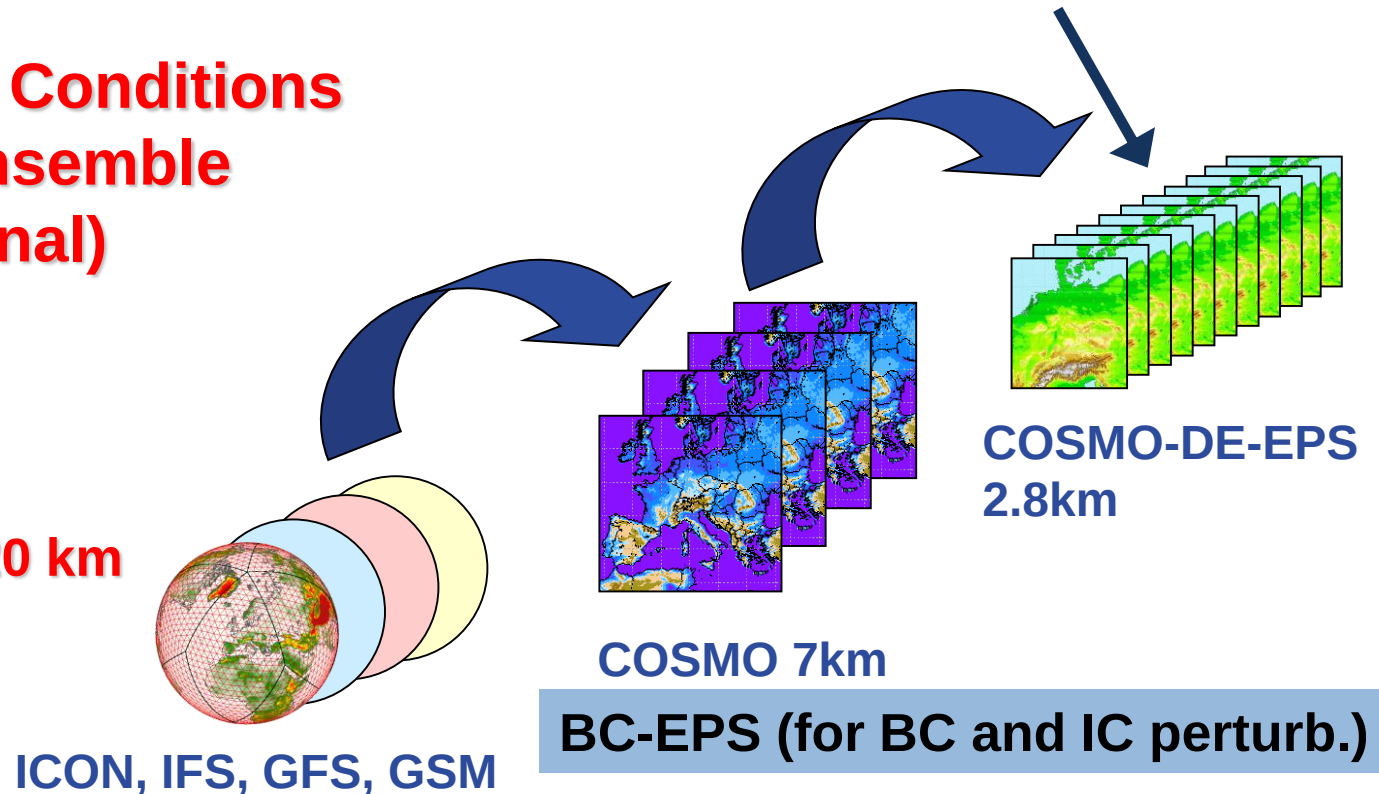


COSMO-DE-EPS operational set-up

- 20 members, grid size: 2.8 km
- 8 starts per day (00, 03, 06,... UTC)
lead time: 0 - 27 hours
0 - 45 hours for 03 UTC

Test: Boundary Conditions from ICON ensemble (pre-operational)

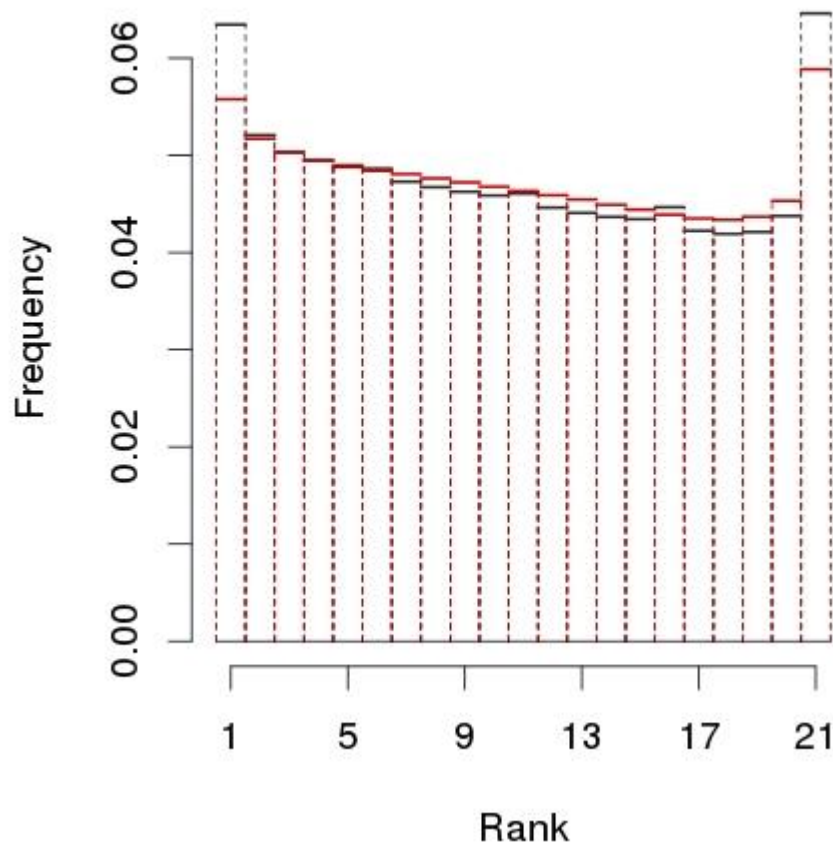
- 40 Member
- Global, 40 km
- ICON-EU Nest, 20 km
- EDA



Rank histogram (hourly precipitation)

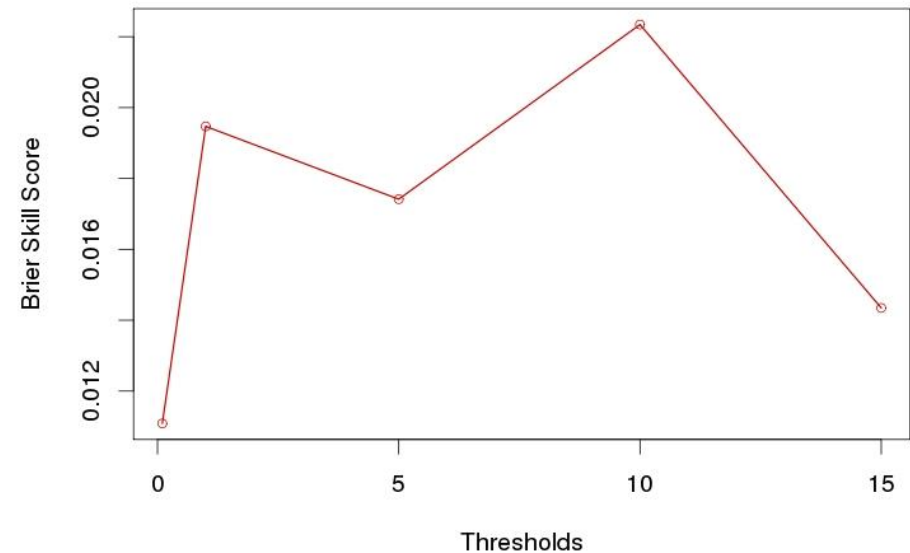
KENDA + ICON-EPS boundary

KENDA + BCEPS boundary



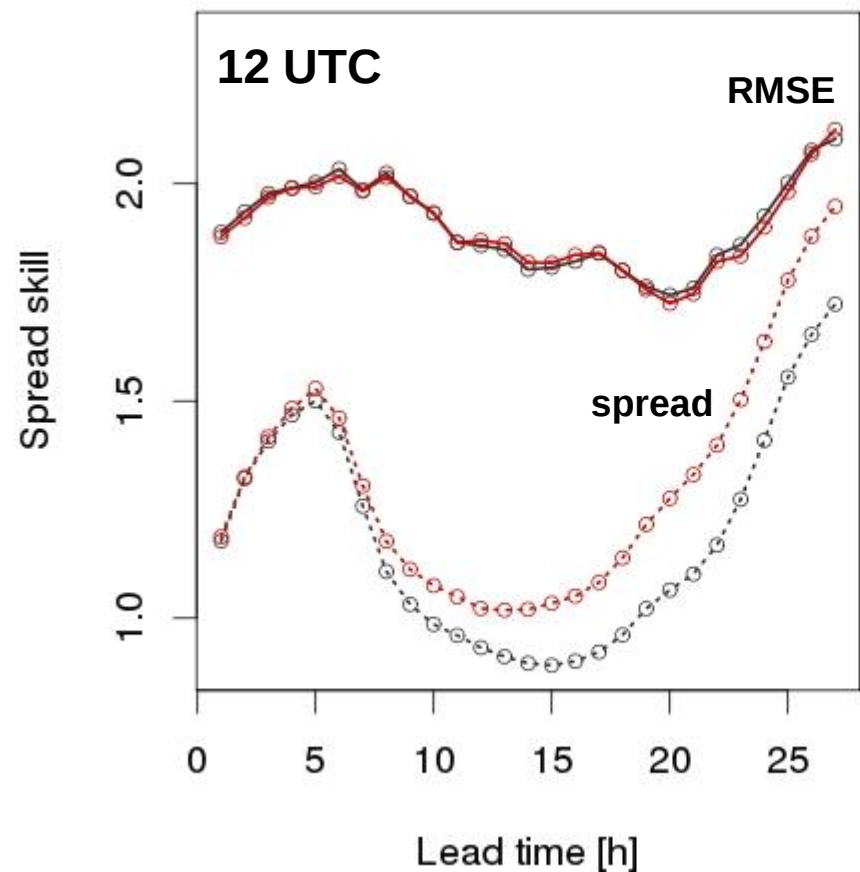
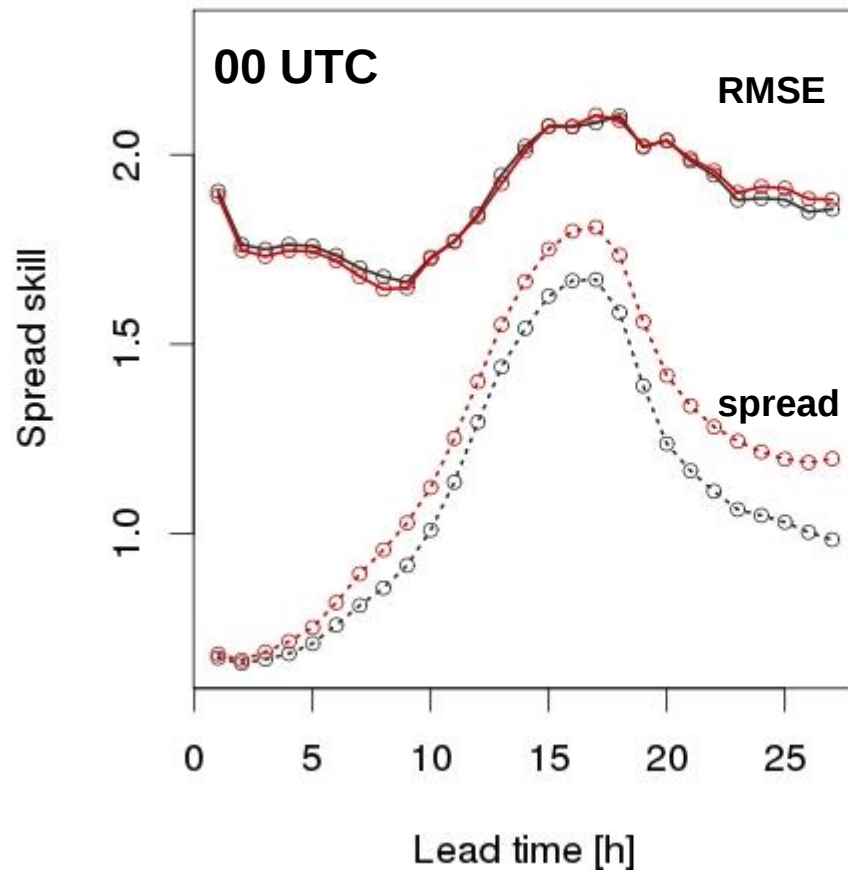
Brier Skill Score (hourly precipitation)

**KENDA + ICON-EPS boundary
(vs KENDA + BCEPS boundary)**



RMSE & spread (wind gusts)

KENDA + ICON-EPS boundary
KENDA + BCEPS boundary

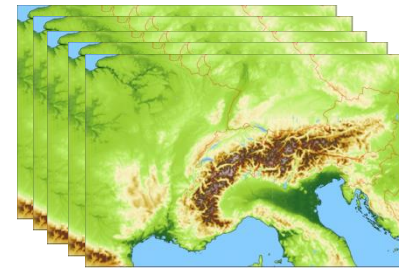


ENSEMBLES:

COSMO-E



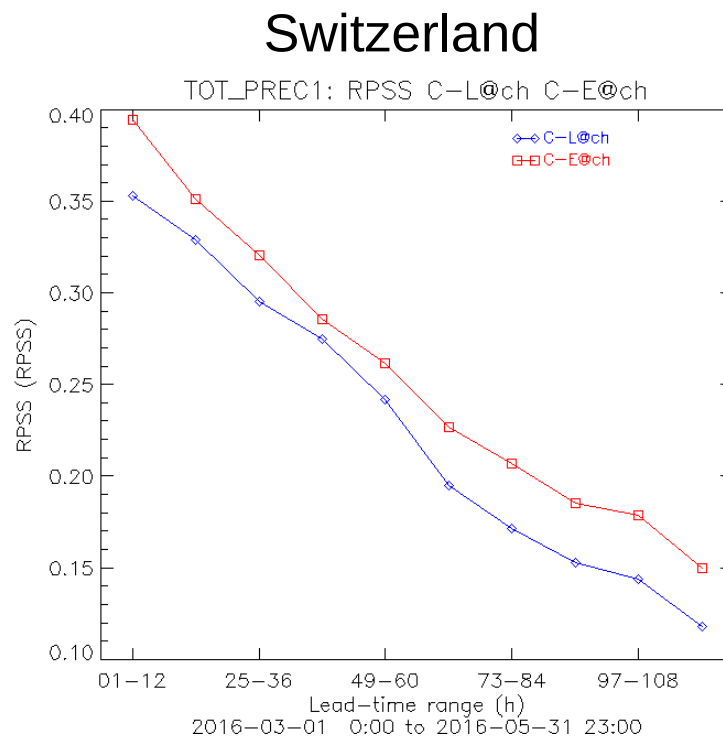
COSMO-E operational setup



- 21 members (control and 20 perturbed runs)
- 2.2 km mesh-size, 60 levels
- two forecasts per day (00 and 12 UTC) up to +120h
- **initial condition** (perturbations): KENDA assimilation cycle
 - KENDA ensemble mean for control
 - KENDA members 1-20 (out of 40)
- **lateral boundary condition** (perturbations): IFS-ENS 18 & 06 UTC (i.e. 6h older LBCs):
 - IFS-ENS control for control
 - IFS-ENS members 1-20 (out of 50)
- **model uncertainty**: SPPT
- COSMO version 5.0+/GPU, single precision

RPSS, 1h precipitation, MAM 2016

COSMO-E
COSMO-LEPS



Thresholds: 0.1,0.2,0.5,1,2,5,10 mm

- COSMO-E shows skill until end of forecast range
- COSMO-E clearly outperforms COSMO-LEPS

SPRED PRIORITY PROJECT

SPREAD/SKILL RELATION SPPT

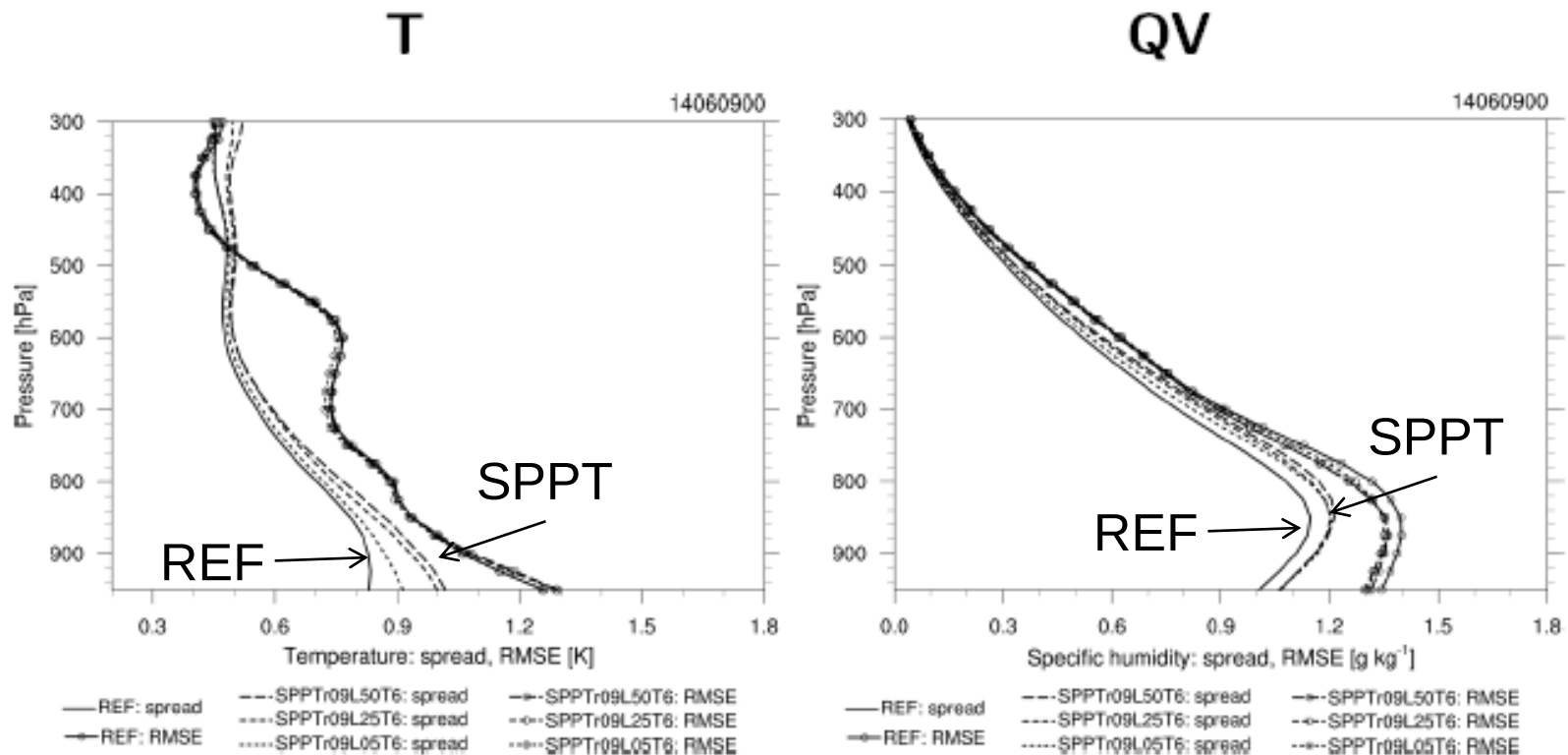
- METEOSWISS (COSMO-E)
- RHM (COSMO-Ru2-EPS)
- IMGW (TLE-MVE)
- ARP AE (COSMO-IT-EPS)



Vertical profiles of spread & RMSE

Convective case (June 2014)

Impact of SPPT



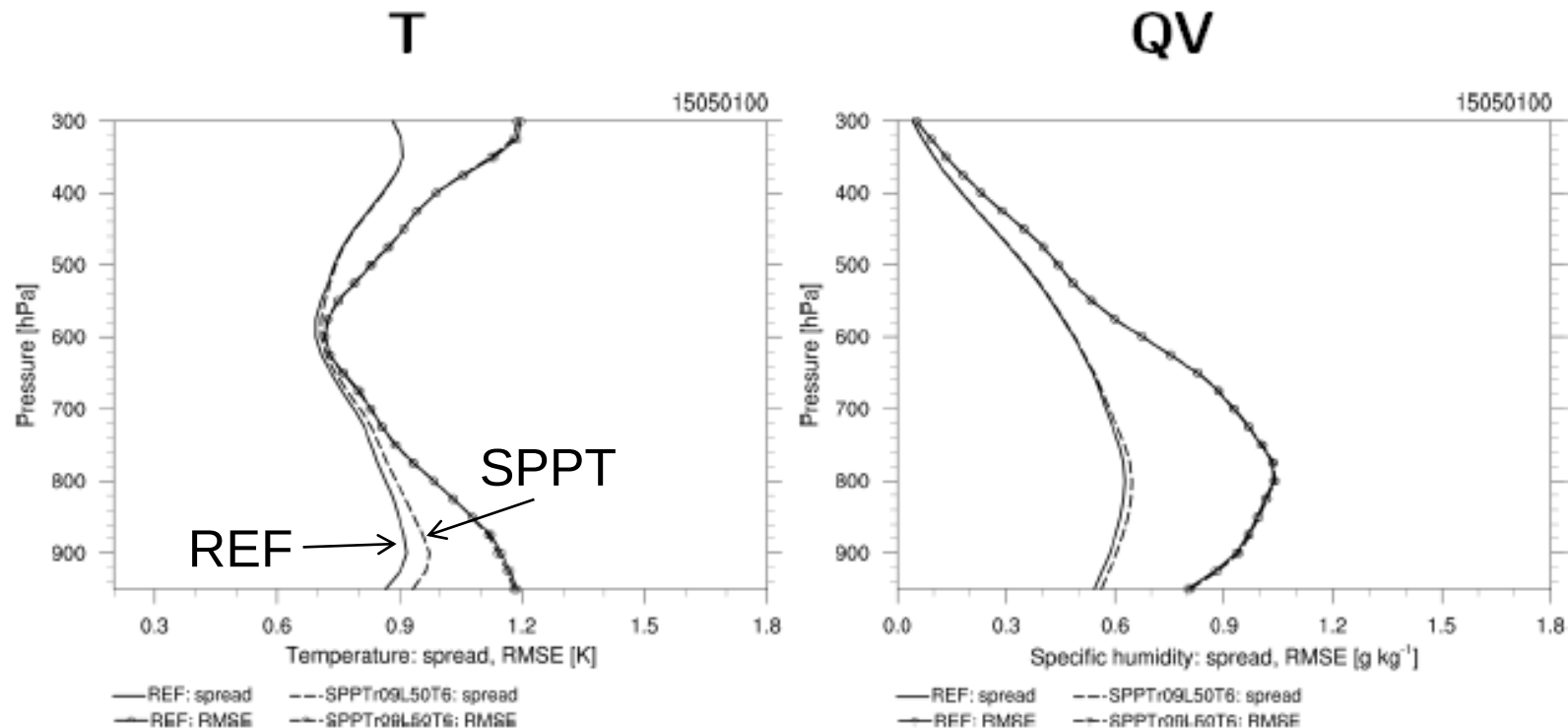
Solid & dashed
With markers
Experiments

Spread
RMSE of ensemble mean
REF and SPPT (5.0°, 2.5°, 0.5°)

Vertical profiles of spread & RMSE

Advective case (May 2015)

Impact of SPPT



Solid & dashed

Spread

With markers

RMSE of ensemble mean

Experiments

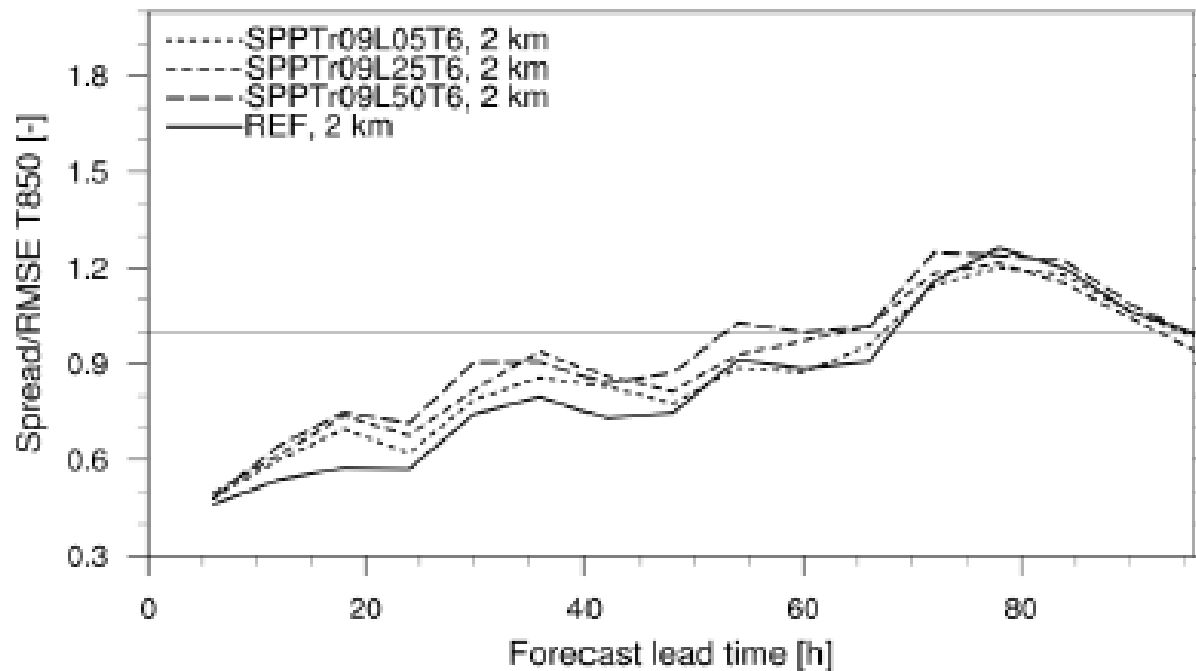
REF and SPPT (5.0°)

Spread & RMSE as function of forecast range

Convective case (June 2014)

Impact of SPPT

SPREAD/RMSE T850

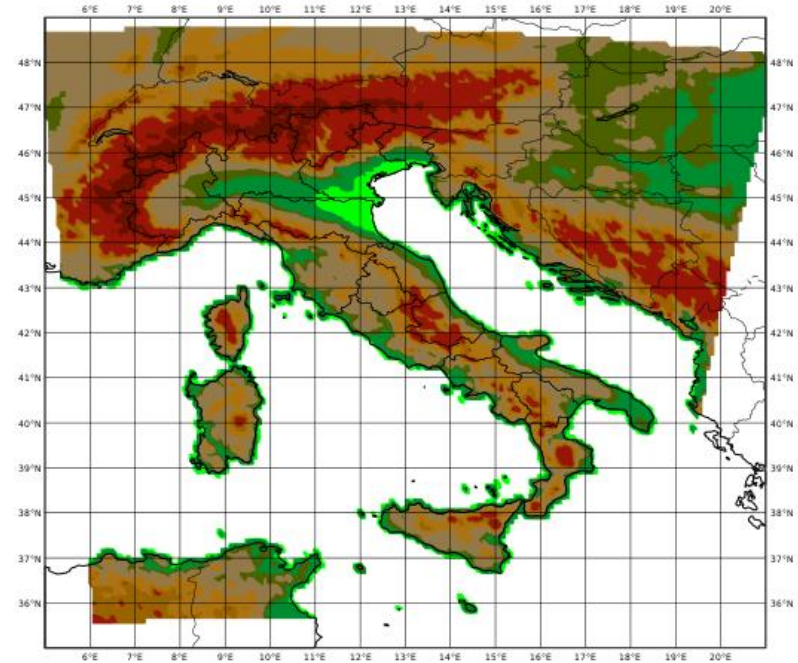


EVALUATION OF ENSEMBLE SPREAD: SAL

- **Aim: assess the impact of physics perturbations on precipitation**
- **What is the perturbation influencing?**
 - **Precipitation intensity**
 - **Precipitation structure**
 - **Localisation of the precipitation**
- ⇒ **Use a spatial verification measure: SAL (Wernli et al 2008)**
- **3 independent components:**
 - **Structure**
 - **Amplitude**
 - **Location**
- **Used here not for verification but for evaluating the similarity between fields, only forecasts**

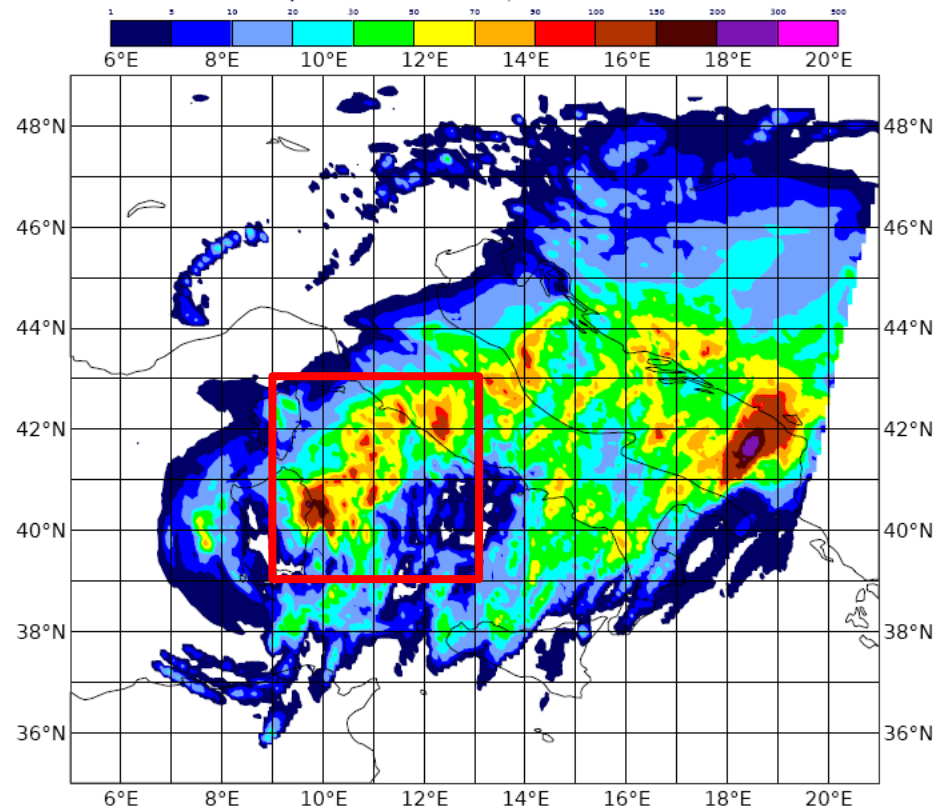
COSMO-IT-EPS

- 2.8 km
- 10 members
- IC/BC from COSMO-ME-EPS
- testing period: October 2015
- 3 set-up for physics perturbations:
 - CTRL: no physics perturbations
 - SPPT: SPPT only
 - SPPT + PP: SPPT + Parameter Perturbation



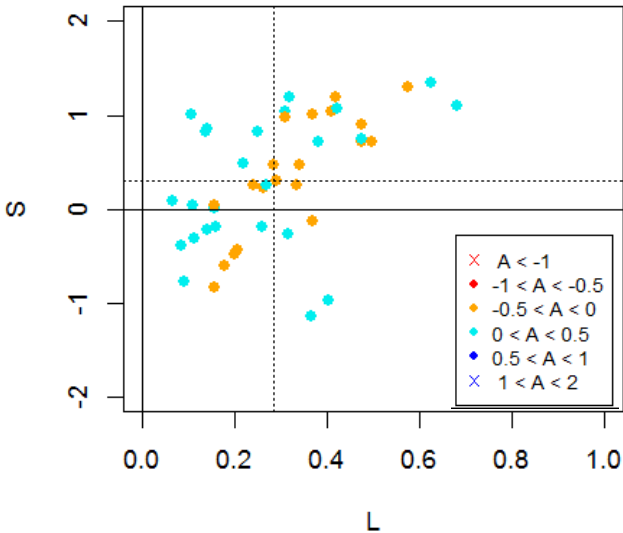
EXAMPLE: INTENSE PRECIPITATION CASE

10/10/15 - 100mm – 125x163

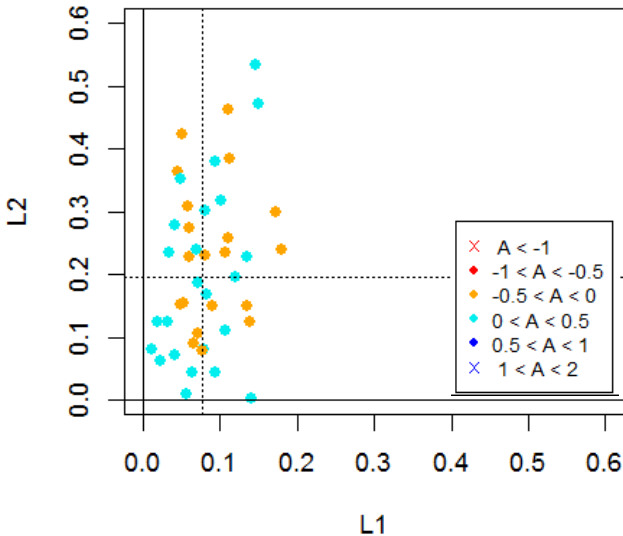


no physics pert

SAL diagram

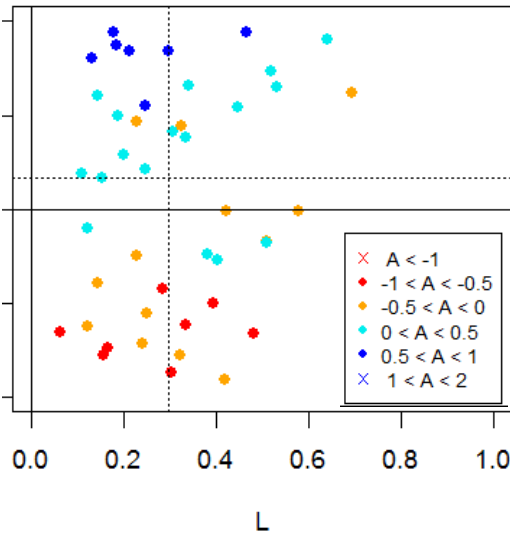


L1-L2 diagram

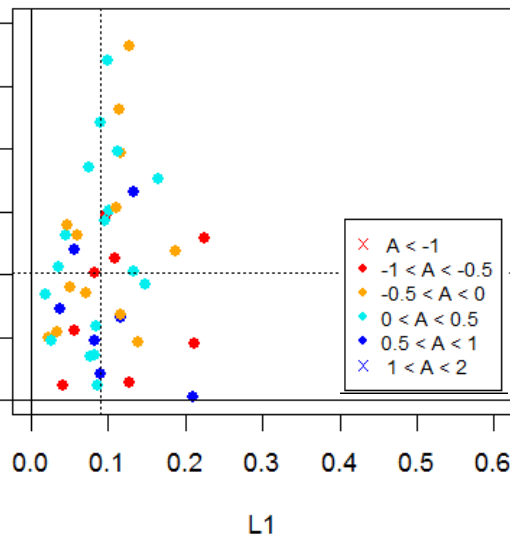


SPPT

SAL diagram

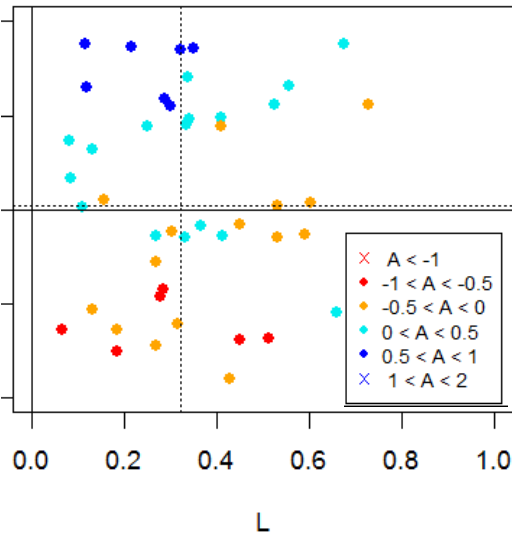


L1-L2 diagram

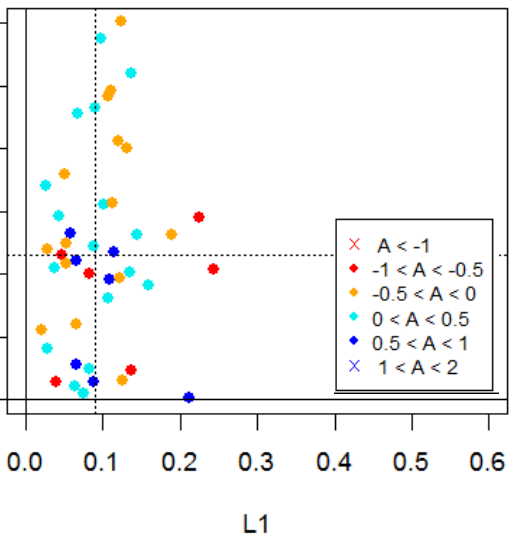


SPPT + PP

SAL diagram



L1-L2 diagram

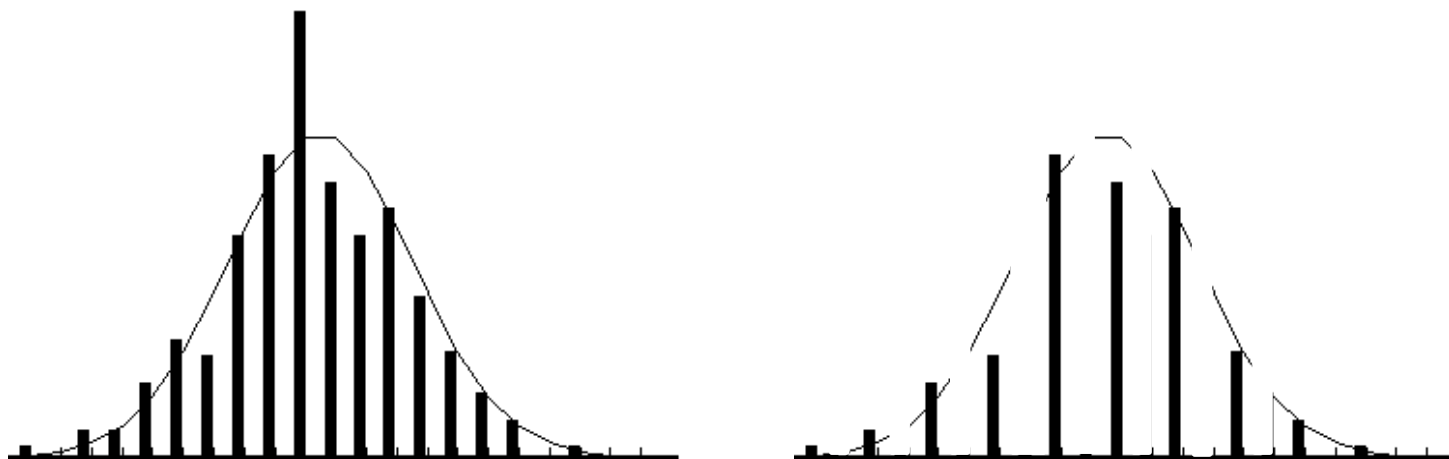


SPRED PRIORITY PROJECT

MEMBER SELECTION FOR CP ENSEMBLES



Goal: Keep the “shape of the PDF”

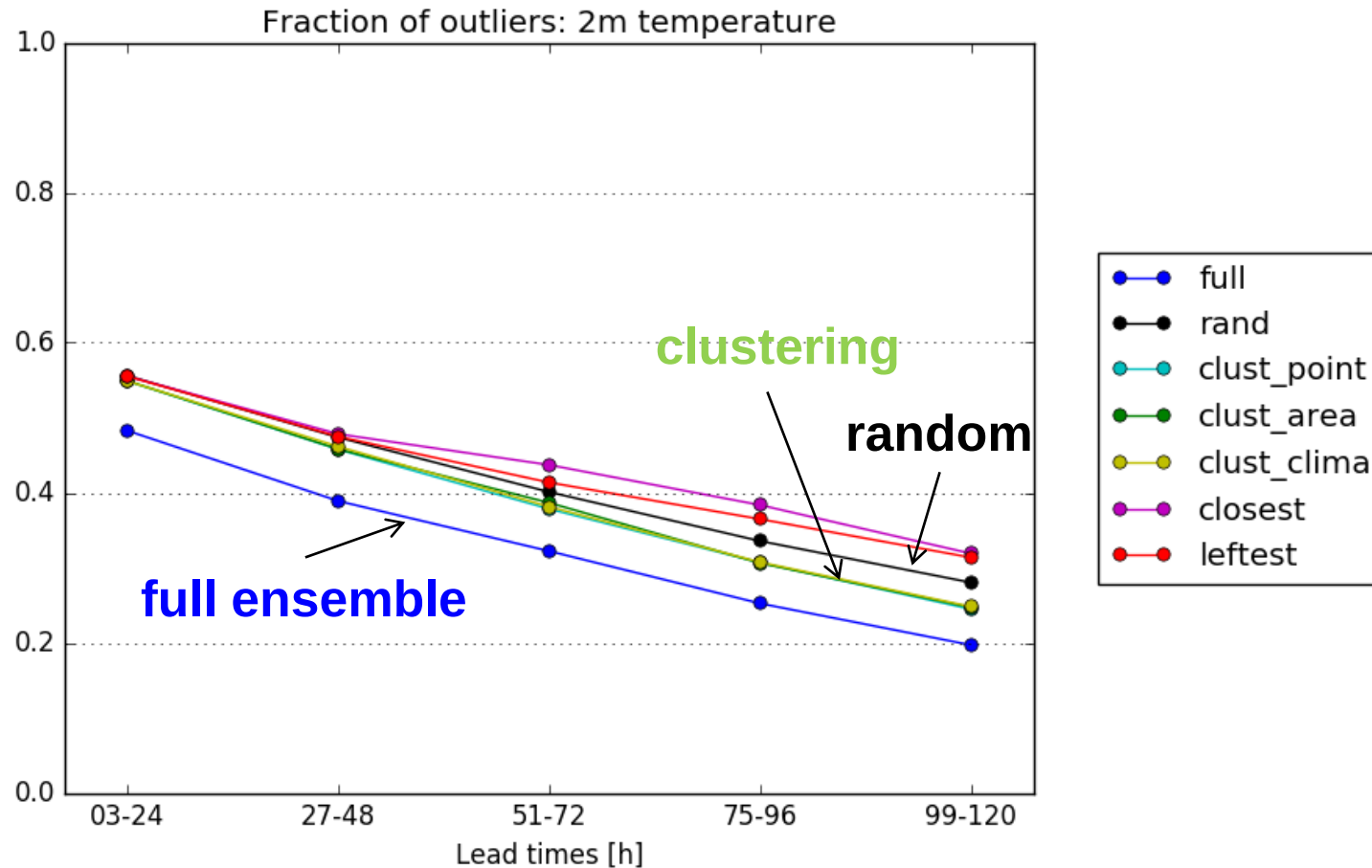


Problem: Multidimensionality (grid-points, variables)

→ reduce phase space and «make» it one-dimensional

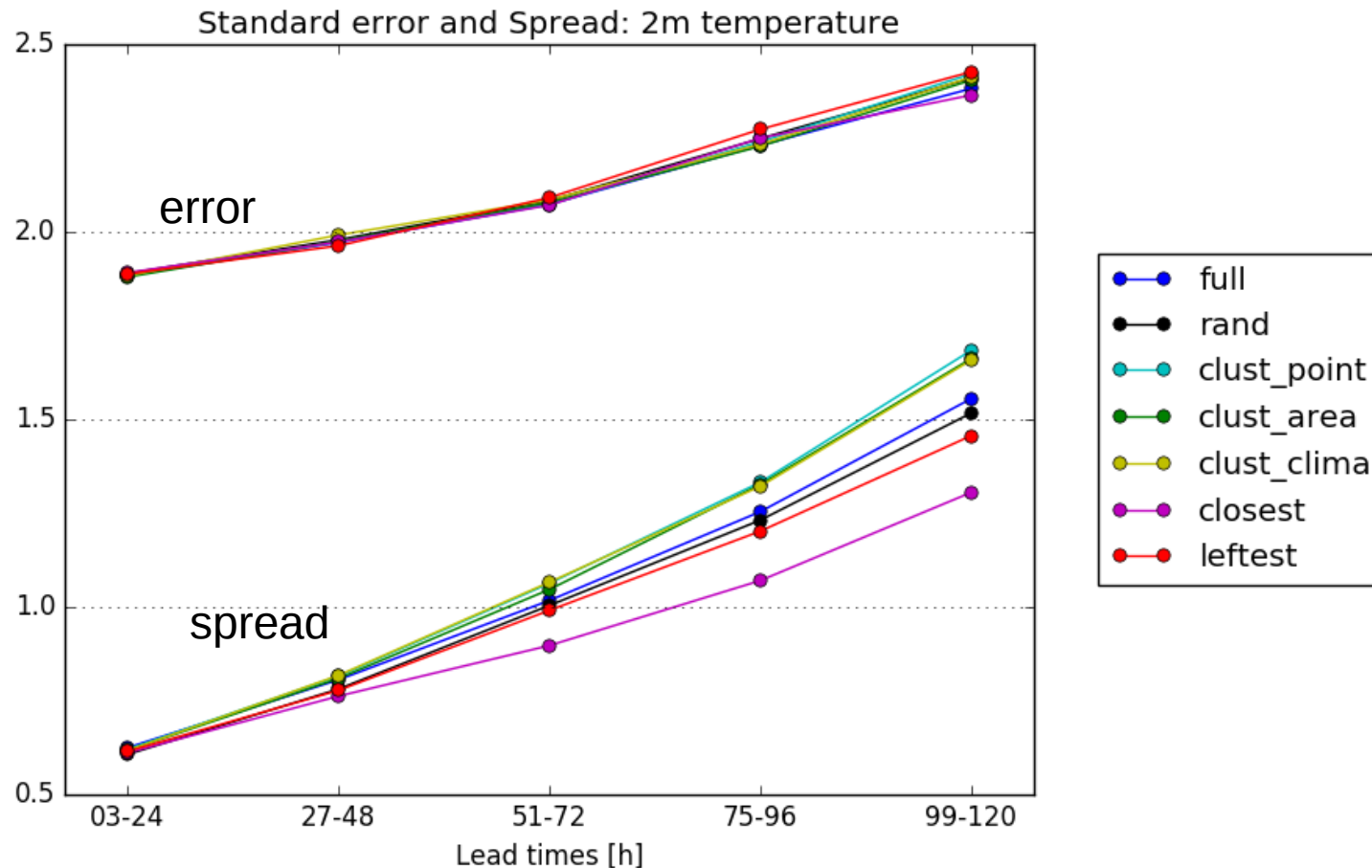
→ similar approach used as in COSMO-LEPS clustering:
3 variables: wind, temperature, humidity on 3 model levels
(~850, 700, 500 hPa)

2m temperature, outliers



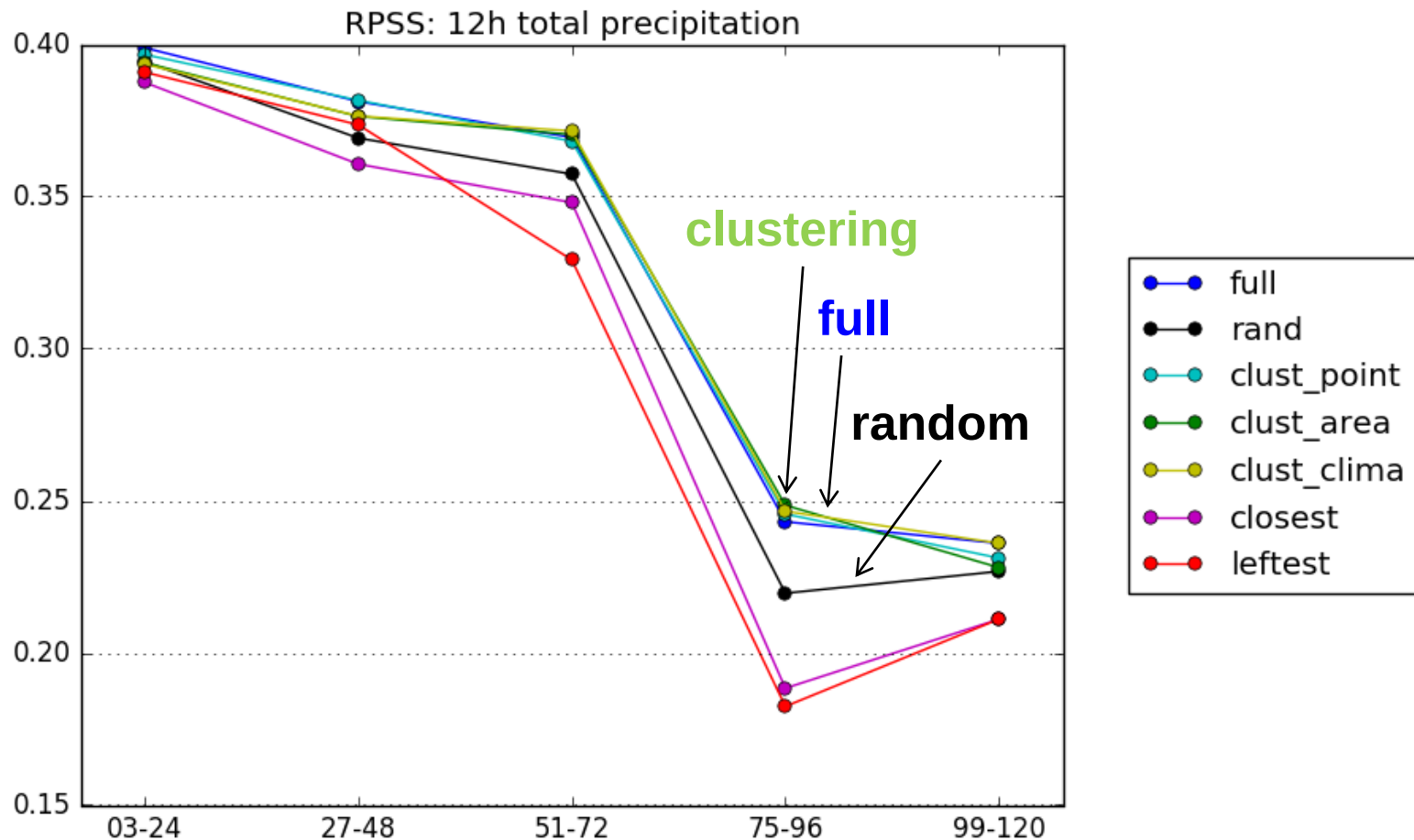
‘full’ best as expected, 3 clustering setups second and almost identical, than ‘rand’, ‘lefttest’, ‘closest’ is worse

2m temperature, spread/error



- 'clust' shows larger spread than 'full'! → tails 'overpopulated'
- 'rand' third, 'closest' clearly worst

12h total precipitation, RPSS



similar results, but clustering as good as 'full' an 'leftest' worst

CONCLUDING REMARKS

- **CP ensembles well established**
- **More knowledge on SPPT impact**
- **Lower boundary perturbations part of the ensemble set-up**
- **Good result on usefulness of member selection for CP ensemble**
- **A report on the spread/skill assessment will be prepared, to summarise what we know (and what we don't know) about the ensemble spread**



THANK YOU!

