### AMPT:

## Additive Model perturbations scaled by Physical Tendency

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- I How model perturbations are generated in DA and EPS?
- Outline of the new scheme termed AMPT: Additive Model perturbations scaled by Physical Tendency
- Testing AMPT in an EPS.

# How model perturbations are generated in DA and EPS?

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DA-specific ways to represent model errors

The most common techniques are:

Multiplicative inflation.

Relaxation of analysis perturbations to the prior ensemble.

Additive inflation.

Disadvantages of these pragmatic approaches:

- Techniques of category (1) provide no additional stochasticity (whereas actual model errors do so).
- Techniques of category (2) are flow independent.
- Both (1) and (2) add perturbations not at sources of model uncertainties.

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### Sources of model errors

- simplifications of model equations
- missing processes
- subgrid-scale processes.

Tackled by physical parameterization schemes  $\Rightarrow$  uncertainty/error in physical tendency

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EPS (and, increasingly, DA):

### Modeling uncertainties in physical parameterizations

Common approaches:

- Multi-physics (ad-hoc, non-stochastic).
- PP (Parameter Perturbations) (ad-hoc, a flavor of multi-physics).
- SPP (Stochastic Parameter Perturbations) (ad-hoc).
- SPPT (Stochastic Perturbations of Physical Tendency) (ad-hoc).
- Intrinsically stochastic physical parameterizations (better justified, promising, but still in their infancy).

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We opted for SPPT because it attempts to do exactly what is needed to represent uncertainty in physical parameterizations: it perturbs the physical tendency.

SPP:

- It accounts only for *parametric* uncertainty (inadequacies in modeling assumptions are not accounted for).
- The parameters may have no counterparts in nature (no objective way to justify the perturbation statistics).

### SPPT: formulation

The SPPT perturbation of the physical tendency in the *i*-th model variable  $P_i$  is

$$\Delta P_i(x, y, \zeta, t) = \epsilon \, \xi(x, y, t) \cdot P_i(x, y, \zeta, t)$$

( $\xi$  is the zero-mean, unit-variance random field,  $\epsilon$  is the magnitude parameter)

#### NB:

The random multiplier  $\epsilon \xi(x, y, t)$  is the same for all model variables and all vertical levels.

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### SPPT: critique

**(**) If at some point  $P_i = 0$ , then the perturbation  $\Delta P_i \propto P_i = 0$ 

(i.e. the assumed model error =0 there).

Can be wrong if, say, in some grid cell, convection is initiated in nature whilst a convective parameterization fails to be activated.

- So The *relative* physical tendency  $\frac{\Delta P_i(x,y,\zeta,t)}{P_i(x,y,\zeta,t)}$  is the same for all model variables at a grid point  $\Rightarrow$  the relative model error is the same for all model variables *i*.
- Similarly, as ξ is constant in the vertical in SPPT, the relative model error is the same for all grid points in the column.
- Moreover, this approximately holds for huge 4D volumes: L=500km (!) and T=6h for SPPT in COSMO.

The SPPT's tacit assumption that errors in different variables everywhere in a LAM domain during hours of forecast time are almost 100% correlated is not realistic.

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# Outline of AMPT

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### AMPT: formulation

From SPPT,

$$\Delta P_i(x, y, \zeta, t) = \epsilon \, \xi(x, y, t) \cdot P_i(x, y, \zeta, t)$$

to

$$\Delta P_i(x, y, \zeta, t) = \epsilon \xi_i(x, y, \zeta, t) \cdot \mathcal{P}_i(x, y, \zeta, t)$$

Differences with SPPT:

- **O** Switch from pointwise physical tendency  $P_i$  to an area-averaged physical tendency  $\mathcal{P}_i$ .
- **2** Specify independent random fields for different model variables  $\xi_i$ .
- Solution Make  $\xi_i$  depend on the vertical coordinate.
- **(4)** Make space and time scales of  $\xi_i$  more realistic for a high-resolution model.

### AMPT in COSMO

- The 4D random pattern  $\xi$  is generated by the Stochastic Pattern Generator (SPG, *Tsyrulnikov and Gayfulin, Meteorol. Zeitschrift, 2017*).
- Perturbed fields:
  - Atmosphere:  $T, u, v, q_v, q_c, q_i$  and hydrostatically balanced p.
  - Soil:  $T_{so}$ ,  $W_{so}$  (multi-layer, 2D random field  $\xi$ ).

### "Gaussian" and "non-Gaussian" fields Example of two unperturbed $T_{soil}$ (left panel) and $W_{soil}$ (right panel) tendency fields Gaussian non-Gaussian



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### Treatment of different model fields

In the soil:

- $T_{so}$ : the area-averaged (scaling) physical tendency  $\mathcal{P}_i$  is computed over the whole LAM domain.
- $W_{so}$ : the scaling physical tendency  $\mathcal{P}_i$  is computed over a small 2D moving window centered at the grid point in question.

In the atmosphere:

- T, u, v are treated like  $T_{so}$
- 2  $q_v, q_c, q_i$  are treated like  $W_{so}$ .

# Testing AMPT in an EPS

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### Experimental setup

•Domain 290\*380 km, centered at Sochi (44N 40E).



•Model: COSMO (version 5.01, single precision), grid spacing 2.2 km, 50 levels. •Ensemble size 10.

•Initial and lateral boundary conditions for ensemble members are taken from COSMO-LEPS adapted for a larger Sochi region (resolution 7 km) — made by the Italian colleagues (special thanks to Andrea Montani).

- •Time period: February March 2014.
- •Verification against synoptic stations.
- •SPG space and time scales:  $L_{\xi}=$  50 km,  $T_{\xi}=$  1 h

### $T_{2m}$ : RMSE of ensemble mean and ensemble spread

Experiment	Model perturbations
NOPERT	None
SPPT	SPPT: atmosphere
AMPT-NOSOIL	AMPT: atmosphere
AMPT-SOIL	$AMPT: \ atmosphere + soil$



### $\Rightarrow$ Spread: big improvement (in *reliability*)

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### Normalized $T_{2m}$ ensemble-mean RMSE reduction



(*RMSE*<sub>NOPERT</sub> – *RMSE*)/*RMSE*<sub>NOPERT</sub> The higher the better. Deterministic verification.

Deterministic verificatio

 $\Rightarrow$  somewhat better

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### $T_{2m}$ : Brier score



The lower the better.

Measures the combined effect of reliability and resolution for the selected threshold.

 $\Rightarrow$  much better.

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### $T_{2m}$ : CRPS

Experiment	Model perturbations
NOPERT	None
SPPT	SPPT: atmosphere
AMPT-NOSOIL	AMPT: atmosphere
AMPT-SOIL	$AMPT: \ atmosphere \ + \ soil$



The lower the better.

Measures the combined effect of reliability and resolution. Integrated over all thresholds.

 $\Rightarrow$  much better

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### $T_{2m}$ : ROC area



The higher the better.

Measures discrimination.

 $\Rightarrow$  much better

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### Precipitation: Brier score



#### The lower the better

 $\Rightarrow \mathsf{Mixed}\ \mathsf{results}$ 

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### Conclusions

- A new model perturbation technique termed AMPT has been developed.
- AMPT aims to address some of the deficiencies of SPPT.
- AMPT generates additive perturbations with the magnitude determined by an area averaged physical tendency.
- AMPT relies on the previously developed 4D Stochastic Pattern Generator (SPG).
- In ensemble prediction experiments:
  - $T, u, v, p_s, q_v, q_c, q_i, T_{so}, W_{so}$  were perturbed.
  - A positive effect from perturbing  $T, u, v, T_{so}, W_{so}$ , mixed effect from perturbing  $q_v, q_c, q_i$ .
  - ► AMPT significantly outperformed SPPT for *T*<sub>2m</sub>, with nearly the same results for precipitation and near-surface wind.